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Gas Dynamics, Lecture 1 (Introduction & Basic Equations)

Practical matters:

- This course:
- Lectures on Monday,
HFML0220; 13.30-15.30;
- Assignment course (werkcollege):
Friday, HGoo.065, 08:30-10:30;
- Lecture Notes and PowerPoint slides on:
www.astro.ru.nl/~achterb/Gasdynamica_2015

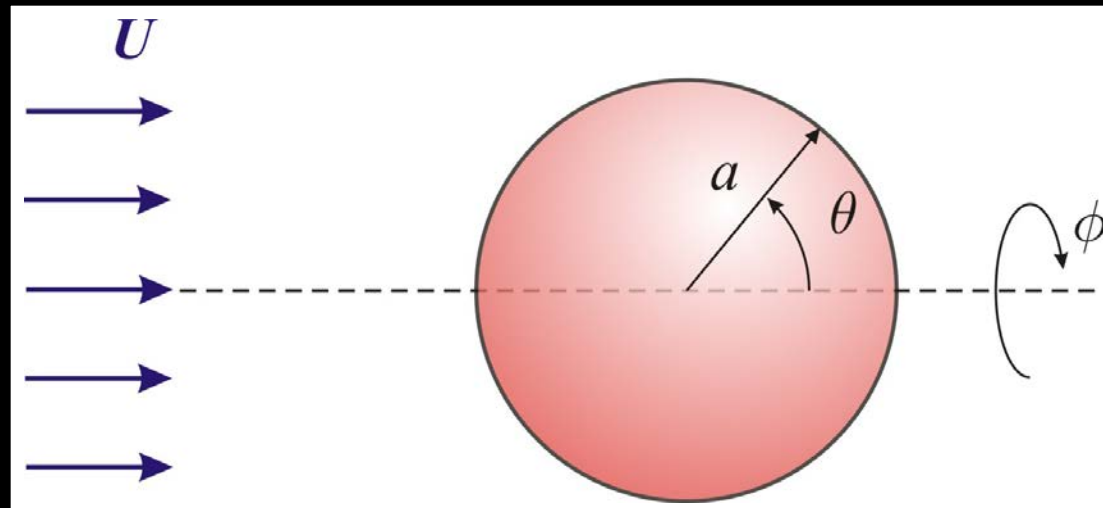
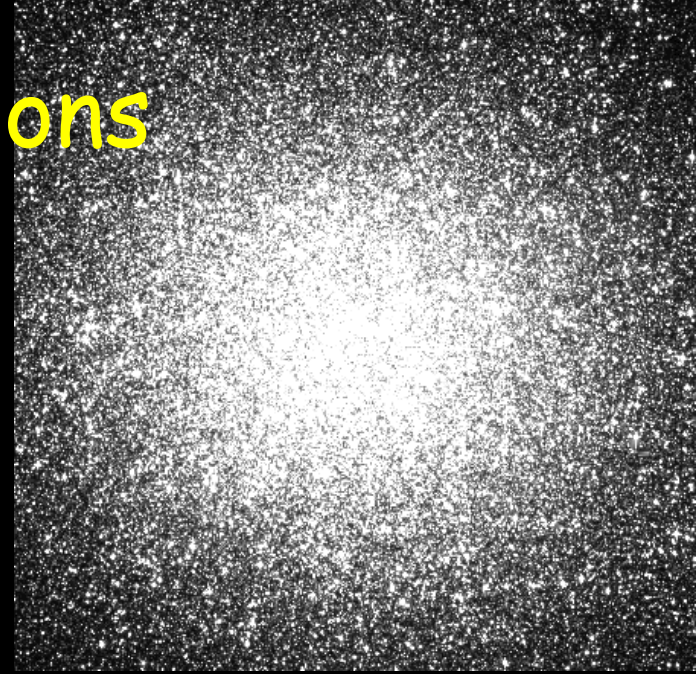
Overview

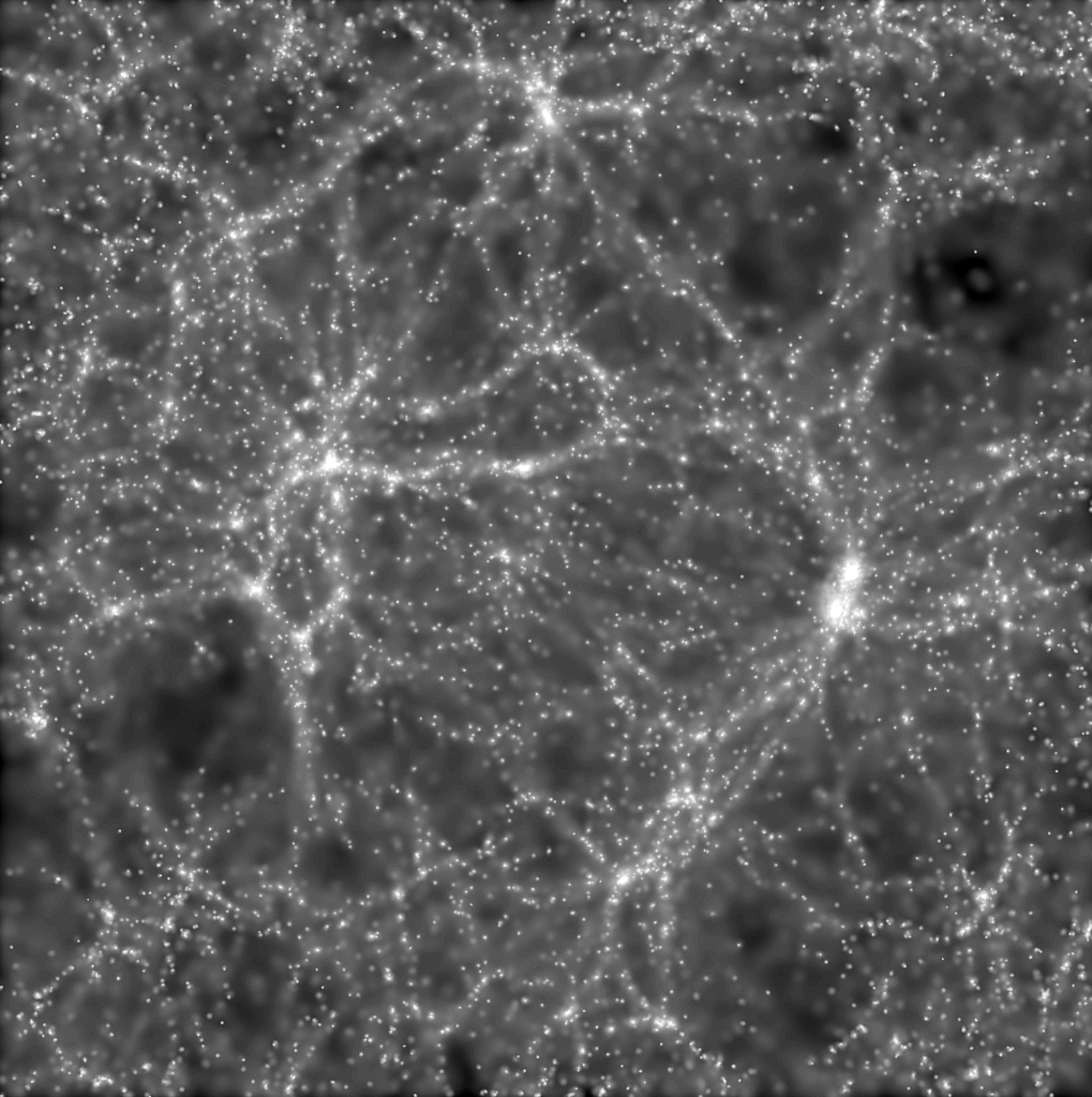
What will we treat during this course?

- Basic equations of gas dynamics
 - Equation of motion
 - Mass conservation
 - Equation of state
- Fundamental processes in a gas
 - Steady Flows
 - Self-gravitating gas
 - Wave phenomena
 - Shocks and Explosions
 - Instabilities: Jeans' Instability

Applications

- Isothermal sphere & Globular Clusters
- Special flows and drag forces
- Solar & Stellar Winds
- Sound waves and surface waves on water
- Shocks
- Point Explosions, Blast waves & Supernova Remnants





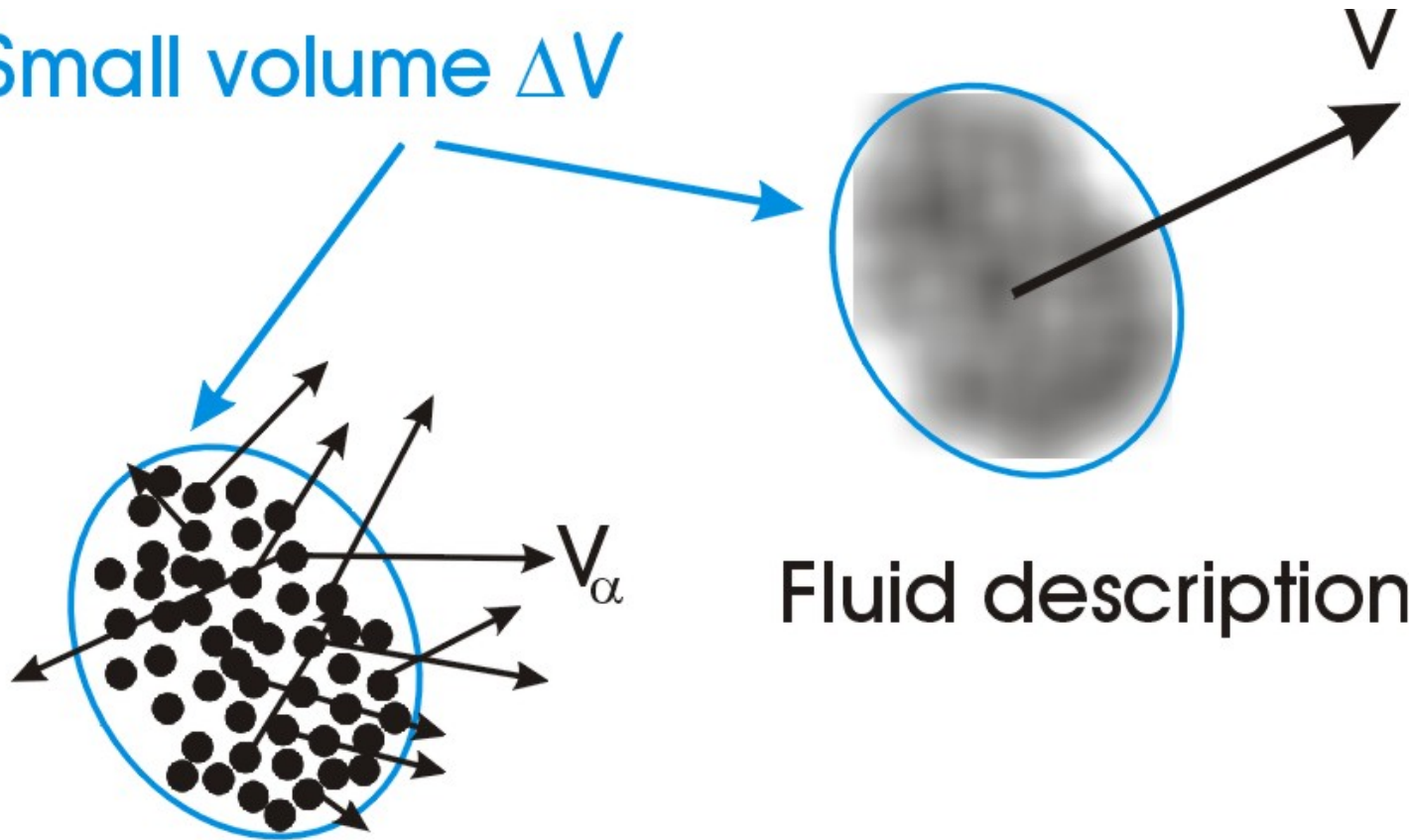
LARGE
SCALE
STRUCTURE

Classical Mechanics vs. Fluid Mechanics

Single-particle (classical) Mechanics	Fluid Mechanics
Deals with <u>single</u> particles with a <u>fixed mass</u>	Deals with a <u>continuum</u> with a <u>variable mass-density</u>
Calculates a <u>single particle trajectory</u>	Calculates a <u>collection of flow lines</u> (flow field) in space
Uses a position <i>vector</i> and velocity <i>vector</i>	Uses a <i>fields</i> : Mass density, velocity field....
Deals only with <u>externally applied</u> forces (e.g. gravity, friction etc)	Deals with <u>internal</u> AND <u>external</u> forces
Is formally linear (so: there is a <u>superposition principle</u> for solutions)	Is intrinsically <u>non-linear</u> <u>No</u> superposition principle in general!

Basic Definitions

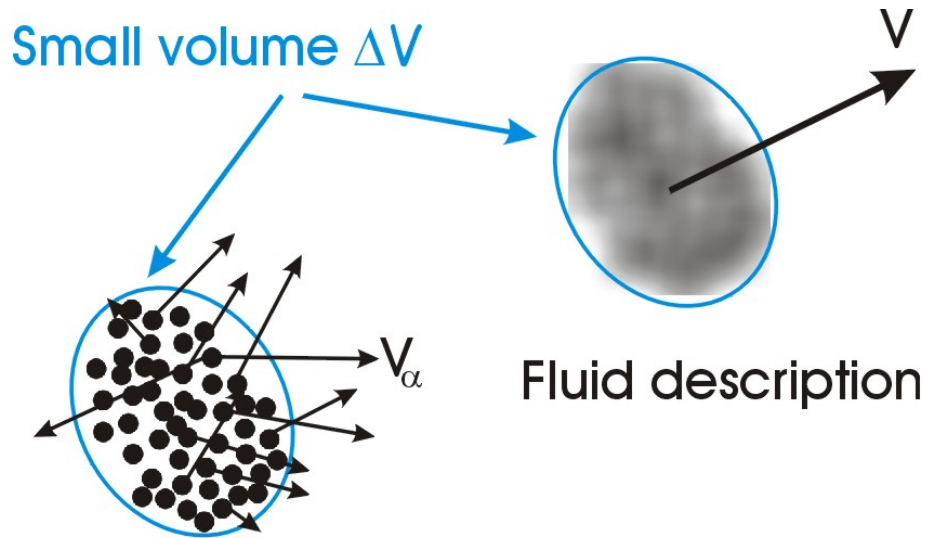
Small volume ΔV



Molecular description

Fluid description

Mass, mass-density and velocity



Mass density ρ :

$$\rho(\mathbf{x}, t) = \lim_{\Delta V \rightarrow 0} \frac{\Delta m}{\Delta V}$$

Molecular description

Mass Δm in volume ΔV

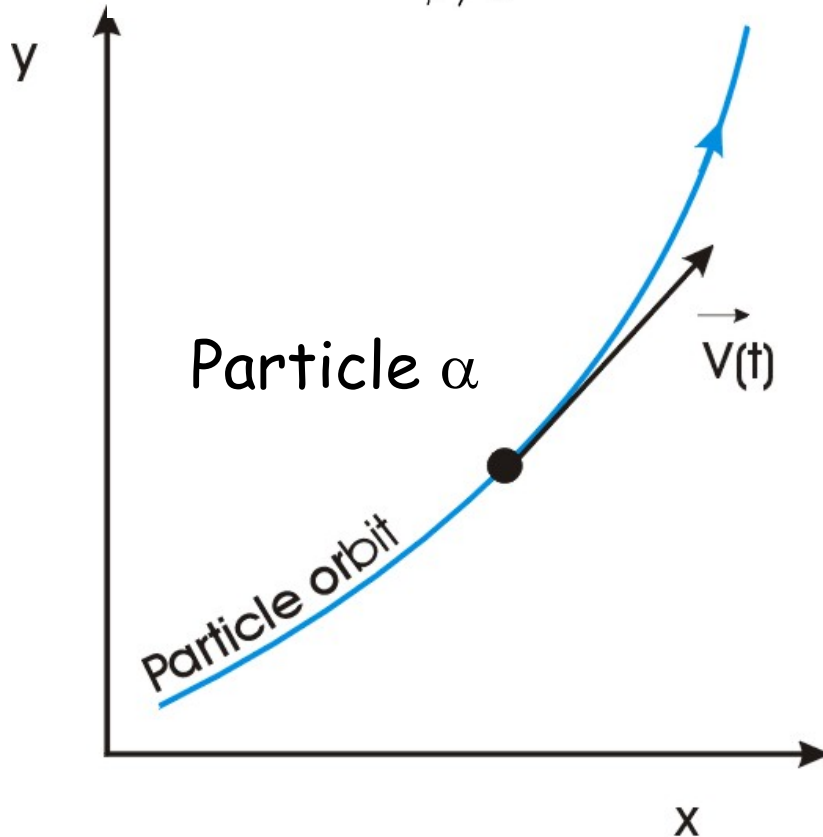
$$\Delta m = \sum_{\mathbf{x}_\alpha \text{ in } \Delta V} m_\alpha$$

Mean velocity $V(\mathbf{x}, t)$
is defined as:

$$\mathbf{V} = \frac{\sum_{\mathbf{x}_\alpha \text{ in } \Delta V} m_\alpha \mathbf{V}_\alpha}{\Delta m}$$

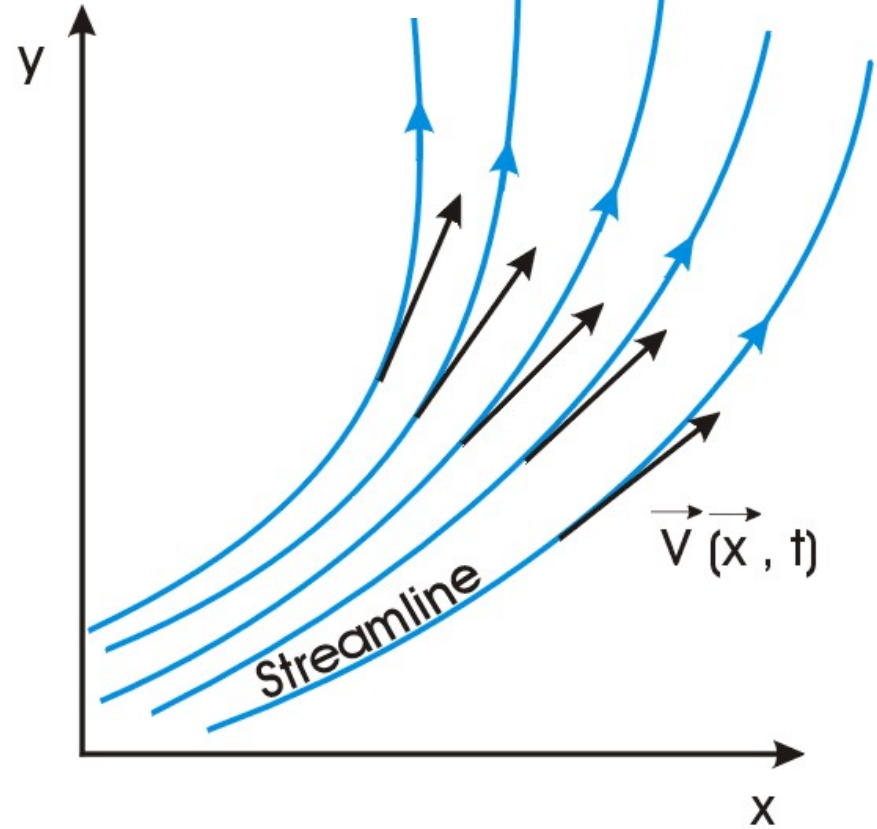
Equation of Motion: from Newton to Navier-Stokes/Euler

$$m_{\alpha} \frac{d\mathbf{V}_{\alpha}}{dt} = \sum_{\beta \neq \alpha} \mathbf{F}_{\alpha\beta}$$



Single-particle dynamics

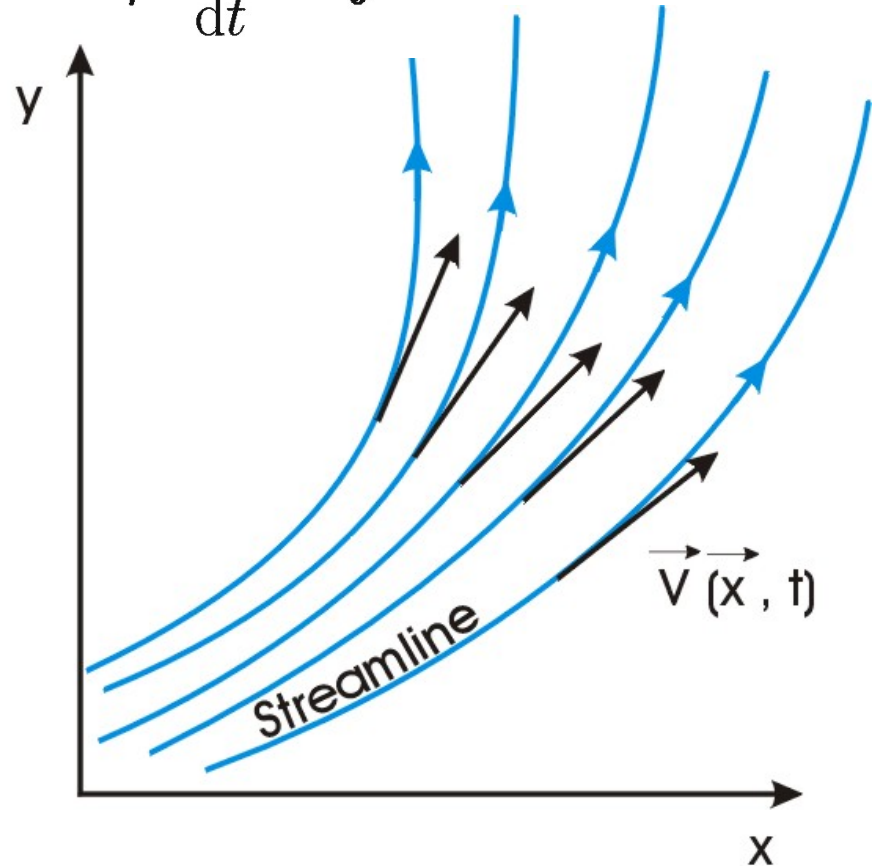
$$\rho \frac{d\mathbf{V}}{dt} = \mathbf{f}$$



Fluid dynamics

Equation of Motion: from Newton to Navier-Stokes/Euler

$$\rho \frac{d\mathbf{V}}{dt} = \mathbf{f}$$



You have to work with a velocity field that depends on position *and* time!

$$\mathbf{V} = (V_x, V_y, V_z) = \mathbf{V}(\mathbf{x}, t)$$

Derivatives, derivatives...

Eulerian change: $\delta Q = Q(\boldsymbol{x}, t + \Delta t) - Q(\boldsymbol{x}, t) \approx \frac{\partial Q}{\partial t} \Delta t$

Derivatives, derivatives...

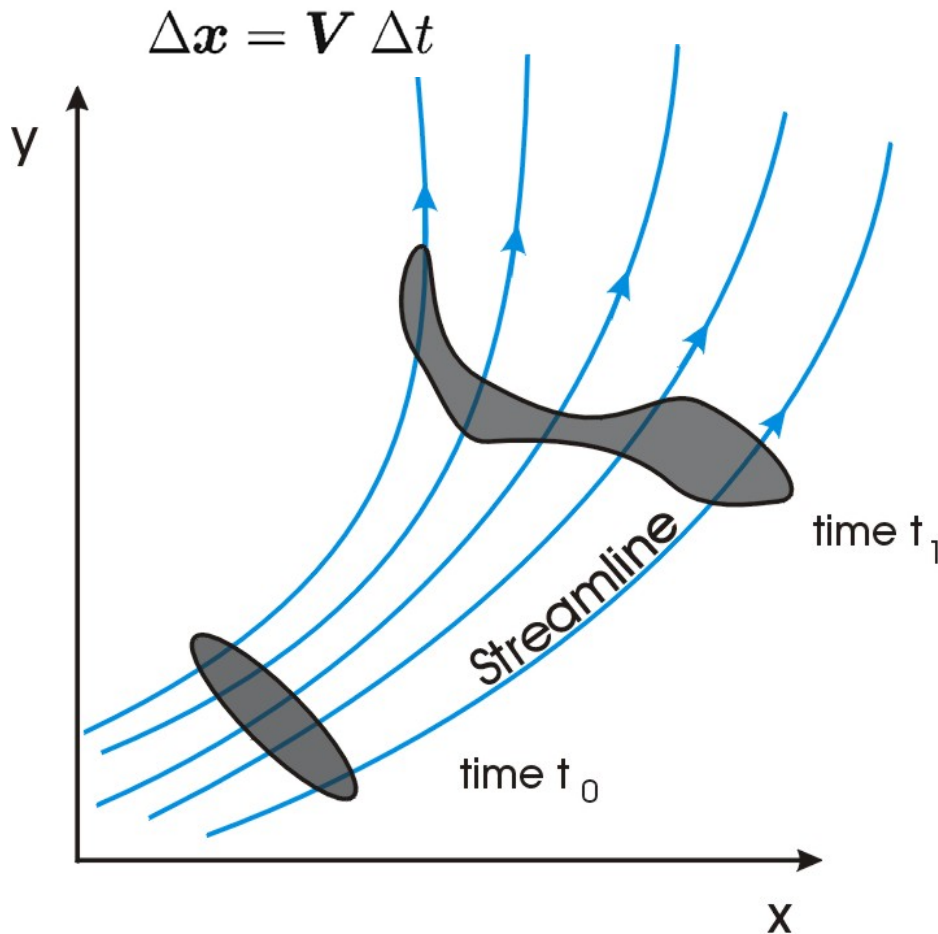
Eulerian change: $\delta Q = Q(\mathbf{x}, t + \Delta t) - Q(\mathbf{x}, t) \approx \frac{\partial Q}{\partial t} \Delta t$
evaluated at a
fixed position

Lagrangian change: $\Delta Q = Q(\mathbf{x} + \Delta \mathbf{x}, t + \Delta t) - Q(\mathbf{x}, t) \approx \frac{dQ}{dt} \Delta t$
evaluated at a
shifting position

Shift along
streamline:

$$\Delta \mathbf{x} = \mathbf{V} \Delta t$$

Comoving derivative d/dt



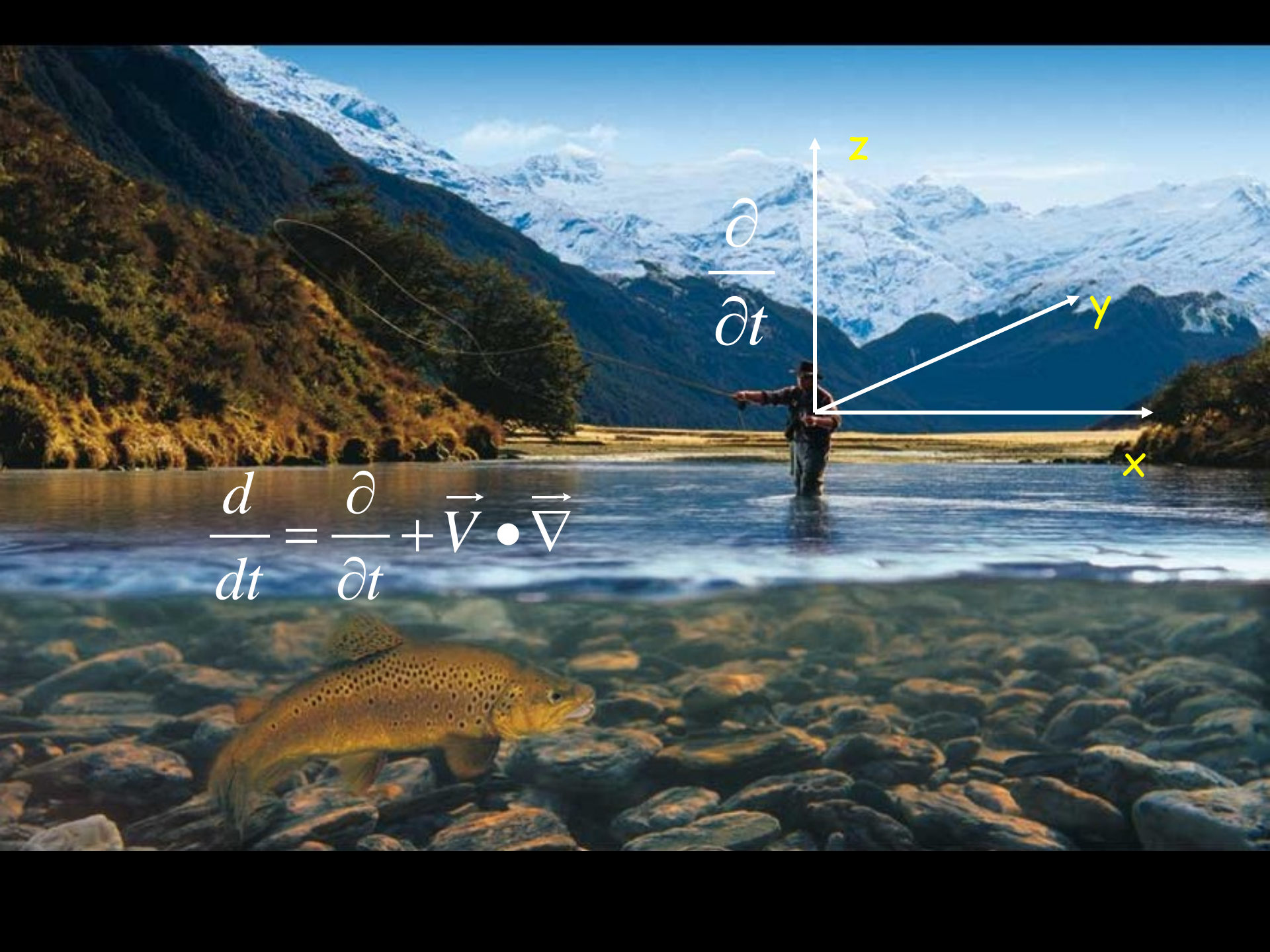
$$\Delta Q = Q(t + \Delta t, \mathbf{x} + \Delta \mathbf{x}) - Q(t, \mathbf{x})$$

$$\approx \frac{\partial Q}{\partial t} \Delta t + (\Delta \mathbf{x} \cdot \nabla) Q$$

$$= \left[\frac{\partial Q}{\partial t} + (\mathbf{V} \cdot \nabla) Q \right] \Delta t$$

$$\equiv \left(\frac{dQ}{dt} \right) \Delta t.$$

$$\frac{d}{dt} = \frac{\partial}{\partial t} + (\mathbf{V} \cdot \nabla)$$



$$\frac{d}{dt} = \frac{\partial}{\partial t} + \vec{V} \cdot \vec{\nabla}$$



Notation: working with the gradient operator

Gradient operator is a 'machine' that converts a scalar into a vector:

$$\nabla = \left(\frac{\partial}{\partial x}, \frac{\partial}{\partial y}, \frac{\partial}{\partial z} \right)$$

For scalar $Q(\mathbf{x}, t)$:

$$\nabla Q = \frac{\partial Q}{\partial x} \hat{e}_x + \frac{\partial Q}{\partial y} \hat{e}_y + \frac{\partial Q}{\partial z} \hat{e}_z$$

Related operators:
turn scalar into scalar,
vector into vector....

$$\Delta \mathbf{x} \cdot \nabla \equiv \Delta x \frac{\partial}{\partial x} + \Delta y \frac{\partial}{\partial y} + \Delta z \frac{\partial}{\partial z}$$

$$\mathbf{V} \cdot \nabla \equiv V_x \frac{\partial}{\partial x} + V_y \frac{\partial}{\partial y} + V_z \frac{\partial}{\partial z}$$

GRADIENT OPERATOR AND VECTOR ANALYSIS (See Appendix A)

scalar into vector: $\mathbf{g} = -\nabla\Phi$

vector into scalar: $\nabla \cdot \mathbf{g} = -4\pi G\rho$

vector into vector: $\nabla \times \mathbf{B} = \frac{4\pi}{c} \mathbf{J}$

tensor into vector: $\nabla \cdot \mathbf{T} = -\mathbf{f}$

Useful relations: $\nabla \cdot (\nabla \times \mathbf{B}) = 0$, $\nabla \times \nabla\Phi = 0$, $\nabla \cdot (\nabla\Phi) = \nabla^2\Phi$

Program for uncovering the basic equations:

1. Define the fluid acceleration and formulate the equation of motion by analogy with single particle dynamics;
2. Identify the forces, such as pressure force;
3. Find equations that describe the response of the other fluid properties (such as: density ρ , pressure P , temperature T) to the flow.

Equation of motion for a fluid:

$$\rho \frac{d\mathbf{V}}{dt} \equiv \rho \left[\frac{\partial \mathbf{V}}{\partial t} + (\mathbf{V} \cdot \nabla) \mathbf{V} \right] = \mathbf{f}$$

Equation of motion for a fluid:

$$\rho \frac{d\mathbf{V}}{dt} \equiv \rho \left[\frac{\partial \mathbf{V}}{\partial t} + (\mathbf{V} \cdot \nabla) \mathbf{V} \right] = \mathbf{f}$$

The acceleration of a fluid element is defined as:

$$\mathbf{a} \equiv \frac{d\mathbf{V}}{dt} = \frac{\partial \mathbf{V}}{\partial t} + (\mathbf{V} \cdot \nabla) \mathbf{V}$$

Equation of motion for a fluid:

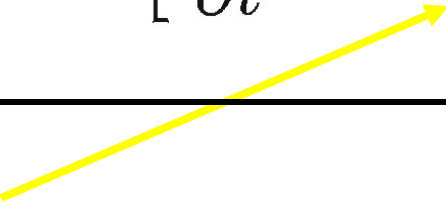
$$\rho \frac{d\mathbf{V}}{dt} \equiv \rho \left[\frac{\partial \mathbf{V}}{\partial t} + (\mathbf{V} \cdot \nabla) \mathbf{V} \right] = \mathbf{f}$$

This equation states:

mass density \times acceleration = force density

note: GENERALLY THERE IS NO
FIXED MASS IN FLUID MECHANICS!

Equation of motion for a fluid:

$$\rho \frac{d\mathbf{V}}{dt} \equiv \rho \left[\frac{\partial \mathbf{V}}{\partial t} + (\mathbf{V} \cdot \nabla) \mathbf{V} \right] = \mathbf{f}$$


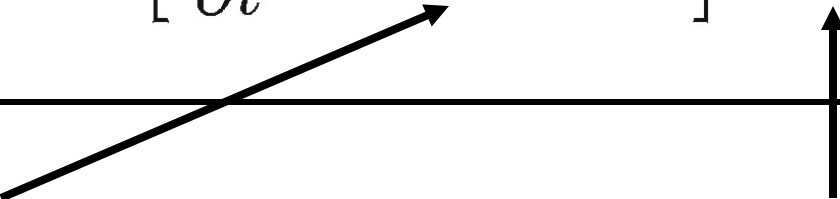
Non-linear term!

Makes it much more difficult
To find 'simple' solutions.

Prize you pay for working with
a velocity-field

$$\mathbf{V} = (V_x, V_y, V_z) = \mathbf{V}(\mathbf{x}, t)$$

Equation of motion for a fluid:

$$\rho \frac{d\mathbf{V}}{dt} \equiv \rho \left[\frac{\partial \mathbf{V}}{\partial t} + (\mathbf{V} \cdot \nabla) \mathbf{V} \right] = \mathbf{f}$$


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Force-density

This force density can be:

- internal:
 - pressure force
 - viscosity (friction)
 - self-gravity
- external
 - For instance: external gravitational force

Pressure force and thermal motions

Split velocities into the
average velocity

$$V(\mathbf{x}, t),$$

and an
isotropically distributed
deviation from average,
the
random velocity:

$$\sigma(\mathbf{x}, t)$$

Individual particle:

$$\mathbf{v}_\alpha = \mathbf{V}(\mathbf{x}, t) + \boldsymbol{\sigma}_\alpha(\mathbf{x}, t) .$$

Average properties of random velocity $\boldsymbol{\sigma}$:

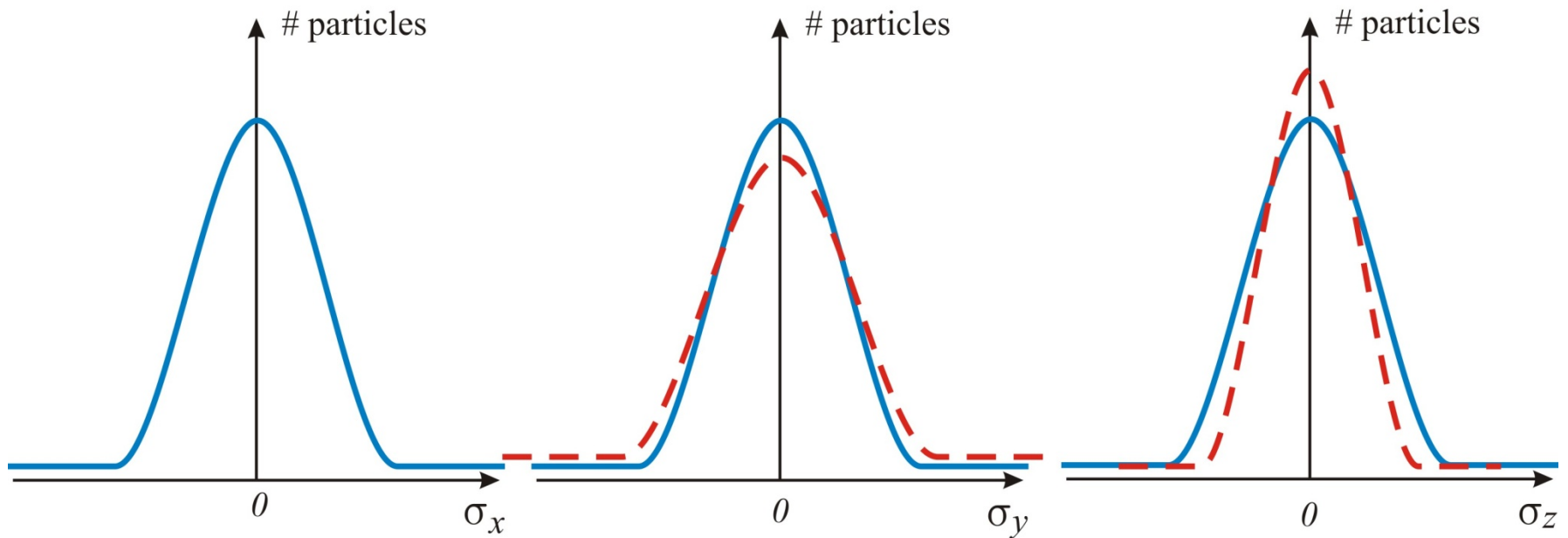
$$\overline{\boldsymbol{\sigma}} = \overline{\mathbf{v}} - \mathbf{V} = \mathbf{0} ;$$

$$\overline{\sigma_x^2} = \overline{\sigma_y^2} = \overline{\sigma_z^2} = \frac{1}{3}\overline{\sigma^2} ,$$

and

$$\overline{\sigma_x \sigma_y} = \overline{\sigma_x \sigma_z} = \overline{\sigma_y \sigma_z} = \dots = 0 .$$

DISTRIBUTION OF RANDOM VELOCITIES ALONG THE THREE COORDINATE AXES



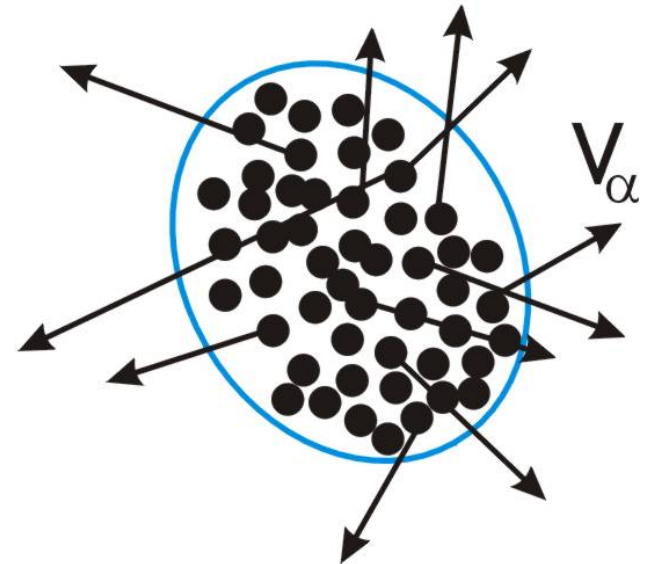
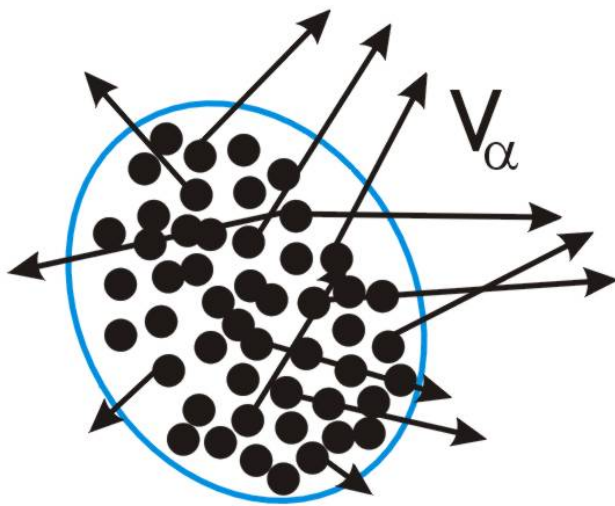
- isotropic case: three distributions identical
- - - anisotropic case: three distributions differ

Mean velocity V

$V=0$



Fluid description



Molecular description

Acceleration of particle α

$$\begin{aligned}\frac{d\mathbf{v}_\alpha}{dt} &= \frac{\partial \mathbf{v}_\alpha}{\partial t} + (\mathbf{v}_\alpha \cdot \nabla) \mathbf{v}_\alpha \\ &= \frac{\partial (\mathbf{V} + \boldsymbol{\sigma}_\alpha)}{\partial t} + ((\mathbf{V} + \boldsymbol{\sigma}_\alpha) \cdot \nabla) (\mathbf{V} + \boldsymbol{\sigma}_\alpha) \\ &= \underbrace{\frac{\partial \mathbf{V}}{\partial t} + (\mathbf{V} \cdot \nabla) \mathbf{V}}_{\text{total derivative mean flow}} + \underbrace{\frac{\partial \boldsymbol{\sigma}_\alpha}{\partial t} + (\mathbf{V} \cdot \nabla) \boldsymbol{\sigma}_\alpha}_{\text{linear in } \boldsymbol{\sigma}} + \underbrace{(\boldsymbol{\sigma}_\alpha \cdot \nabla) \boldsymbol{\sigma}_\alpha}_{\text{quadratic in } \boldsymbol{\sigma}}\end{aligned}$$

Acceleration of particle α (II)

Effect of average over many particles in small volume:

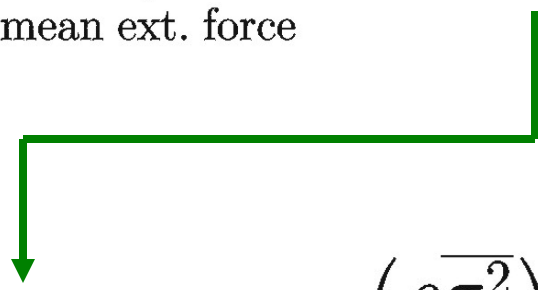
$$\begin{aligned}\overline{\frac{d\mathbf{v}}{dt}} &= \overline{\frac{\partial \mathbf{v}}{\partial t} + (\mathbf{v} \cdot \nabla) \mathbf{v}} \\ &= \underbrace{\frac{\partial \mathbf{V}}{\partial t} + (\mathbf{V} \cdot \nabla) \mathbf{V}}_{\text{total derivative mean flow}} + \underbrace{\left(\frac{\partial}{\partial t} + (\mathbf{V} \cdot \nabla) \right) \overline{\boldsymbol{\sigma}}}_{\text{vanishes: } \overline{\boldsymbol{\sigma}}=0!} + \underbrace{\overline{(\boldsymbol{\sigma} \cdot \nabla) \boldsymbol{\sigma}}}_{\text{remains: quadratic in } \boldsymbol{\sigma}}\end{aligned}$$

Average equation of motion:

$$\rho \overline{\frac{d\mathbf{v}}{dt}} = \overline{\mathbf{f}}$$

$$\rho \left(\frac{\partial \mathbf{V}}{\partial t} + (\mathbf{V} \cdot \nabla) \mathbf{V} \right) = \underbrace{\overline{\mathbf{f}}}_{\text{mean ext. force}} - \rho \overline{(\boldsymbol{\sigma} \cdot \nabla) \boldsymbol{\sigma}}$$

For isotropic fluid:

$$\rho \overline{(\boldsymbol{\sigma} \cdot \nabla) \boldsymbol{\sigma}} = \nabla \left(\frac{\overline{\rho \sigma^2}}{3} \right) \equiv \nabla P$$


Some tensor algebra:

Vector

$$\mathbf{A} \equiv A_i \mathbf{e}_i = A_x \mathbf{e}_1 + A_y \mathbf{e}_2 + A_z \mathbf{e}_3 = \begin{pmatrix} A_x \\ A_y \\ A_z \end{pmatrix}$$

Three notations for the same animal!

Some tensor algebra:

the divergence of a vector in cartesian
(x, y, z) coordinates

Vector



Scalar

$$\mathbf{A} \equiv A_i \mathbf{e}_i = A_x \mathbf{e}_1 + A_y \mathbf{e}_2 + A_z \mathbf{e}_3 = \begin{pmatrix} A_x \\ A_y \\ A_z \end{pmatrix}$$

$$\nabla \cdot \mathbf{A} = \frac{\partial A_i}{\partial x_i} = \frac{\partial A_x}{\partial x} + \frac{\partial A_y}{\partial y} + \frac{\partial A_z}{\partial z}$$

Rank 2 Tensor

Rank 2
tensor

$$\mathbf{T} = T_{ij} \mathbf{e}_i \otimes \mathbf{e}_j \implies \begin{pmatrix} T_{xx} & T_{xy} & T_{xz} \\ T_{yx} & T_{yy} & T_{yz} \\ T_{zx} & T_{zy} & T_{zz} \end{pmatrix}$$

Rank 2 Tensor and Tensor Divergence

Rank 2
tensor \mathbf{T}



Vector

$$\mathbf{T} = T_{ij} \mathbf{e}_i \otimes \mathbf{e}_j = \begin{pmatrix} T_{xx} & T_{xy} & T_{xz} \\ T_{yx} & T_{yy} & T_{yz} \\ T_{zx} & T_{zy} & T_{zz} \end{pmatrix}$$

$$\nabla \cdot \mathbf{T} = \left(\frac{\partial T_{ij}}{\partial x_i} \right) \mathbf{e}_j = \begin{pmatrix} \frac{\partial T_{xx}}{\partial x} + \frac{\partial T_{yx}}{\partial y} + \frac{\partial T_{zx}}{\partial z} \\ \frac{\partial T_{xy}}{\partial x} + \frac{\partial T_{yy}}{\partial y} + \frac{\partial T_{zy}}{\partial z} \\ \frac{\partial T_{xz}}{\partial x} + \frac{\partial T_{yz}}{\partial y} + \frac{\partial T_{zz}}{\partial z} \end{pmatrix}$$

Special case:

Dyadic Tensor = Direct Product of two Vectors

$$\mathbf{A} \otimes \mathbf{B} \equiv A_i B_j \mathbf{e}_i \otimes \mathbf{e}_j = \begin{pmatrix} A_x B_x & A_x B_y & A_x B_z \\ A_y B_x & A_y B_y & A_y B_z \\ A_z B_x & A_z B_y & A_z B_z \end{pmatrix}$$

$$\nabla \cdot (\mathbf{A} \otimes \mathbf{B}) = (\nabla \cdot \mathbf{A}) \mathbf{B} + (\mathbf{A} \cdot \nabla) \mathbf{B}$$

This is the product rule for differentiation!

Application: Pressure Force (I)

Tensor divergence: $(\rho \boldsymbol{\sigma} \cdot \nabla) \boldsymbol{\sigma} = \nabla \cdot (\rho \boldsymbol{\sigma} \otimes \boldsymbol{\sigma}) - (\nabla \cdot (\rho \boldsymbol{\sigma})) \boldsymbol{\sigma}$

Isotropy of the random velocities:

$$\rho \overline{(\boldsymbol{\sigma} \cdot \nabla) \boldsymbol{\sigma}} = \nabla \cdot (\rho \overline{\boldsymbol{\sigma} \otimes \boldsymbol{\sigma}})$$

Second term = scalar x vector!

This must vanish upon averaging!!

Application: Pressure Force (II)

Isotropy of the random velocities

$$\rho \overline{(\boldsymbol{\sigma} \cdot \nabla) \boldsymbol{\sigma}} = \nabla \cdot (\rho \overline{\boldsymbol{\sigma} \otimes \boldsymbol{\sigma}})$$

$$\overline{\sigma_i \sigma_j} = \frac{1}{3} \overline{\sigma^2} \delta_{ij} = \begin{cases} \frac{1}{3} \overline{\sigma^2} & \text{when } i = j \\ 0 & \text{when } i \neq j \end{cases}$$

$$\rho \overline{\boldsymbol{\sigma} \otimes \boldsymbol{\sigma}} = \rho \begin{pmatrix} \frac{1}{3} \overline{\sigma^2} & 0 & 0 \\ 0 & \frac{1}{3} \overline{\sigma^2} & 0 \\ 0 & 0 & \frac{1}{3} \overline{\sigma^2} \end{pmatrix} = \frac{\rho \overline{\sigma^2}}{3} \mathbf{I}$$

Diagonal Pressure Tensor

Pressure force, conclusion:

$$\rho \overline{(\boldsymbol{\sigma} \cdot \nabla) \boldsymbol{\sigma}} = \nabla \cdot (\rho \overline{\boldsymbol{\sigma} \otimes \boldsymbol{\sigma}}) = \nabla \left(\frac{\rho \overline{\sigma^2}}{3} \right) \equiv \nabla P$$

Equation of motion for frictionless ('ideal') fluid:

$$\rho \left(\frac{\partial \mathbf{V}}{\partial t} + (\mathbf{V} \cdot \nabla) \mathbf{V} \right) = -\nabla P + \text{other (external) forces}$$

$$P(\mathbf{x}, t) \equiv \frac{1}{3} \rho \overline{\sigma^2}$$

Summary:

- We know how to interpret the time-derivative d/dt ;
- We know what the equation of motion looks like;
- We know where the pressure force comes from (thermal motions), and how it looks: $f = -\nabla P$.
- **We still need:**
 - A way to link the pressure to density and temperature: $P = P(\rho, T)$;
 - A way to calculate how the density ρ of the fluid changes.