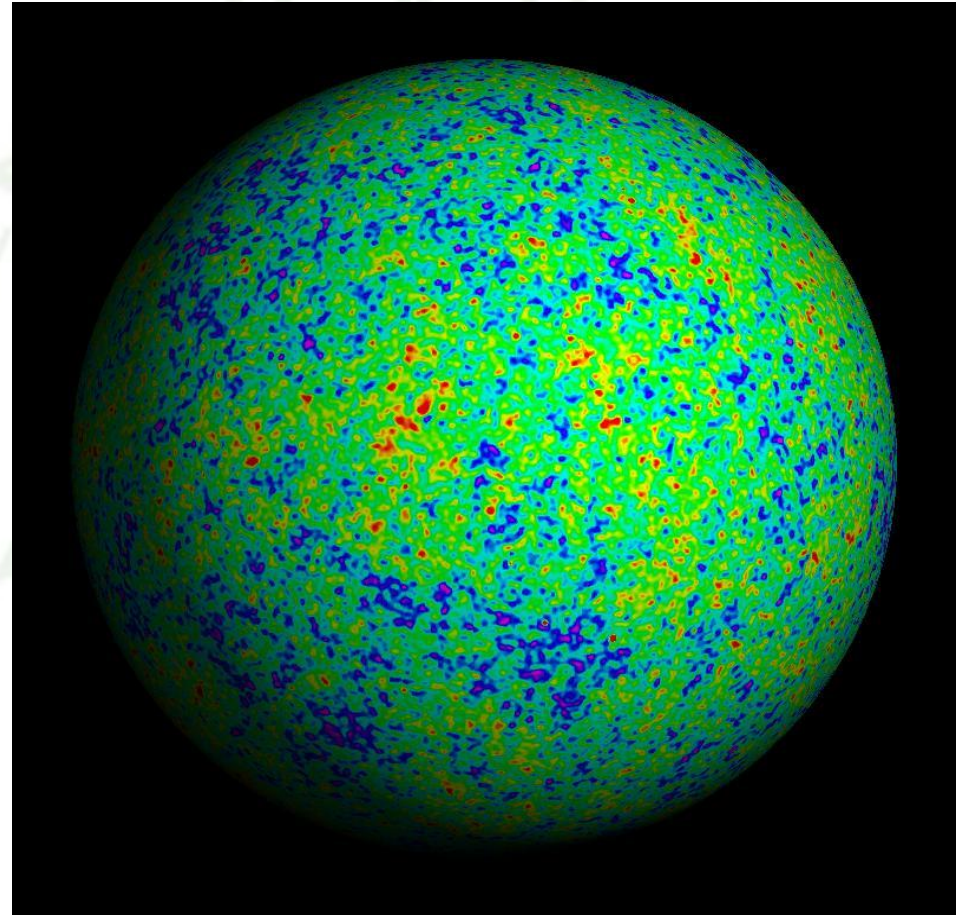


# HOVO cursus Kosmologie

Voorjaar 2011

prof.dr. Paul Groot  
dr. Gijs Nelemans

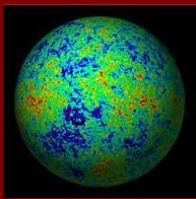


Afdeling Sterrenkunde,  
Radboud Universiteit Nijmegen

# HOVO cursus Kosmologie

## Overzicht van de cursus:

- 17/1 Groot Historische inleiding + afstandsladder
- 24/1 *Nelemans* *Observationele kosmologie, CMB*
- 31/1 Nelemans Friedmann vgl., Nucleosynthese
- 7/2 Nelemans Donkere materie en Structuurvorming
- 14/2 Nelemans Supernovae, donkere energie
- 21/2 Groot Weak lensing, EUCLID, toekomst



# De wet van Hubble

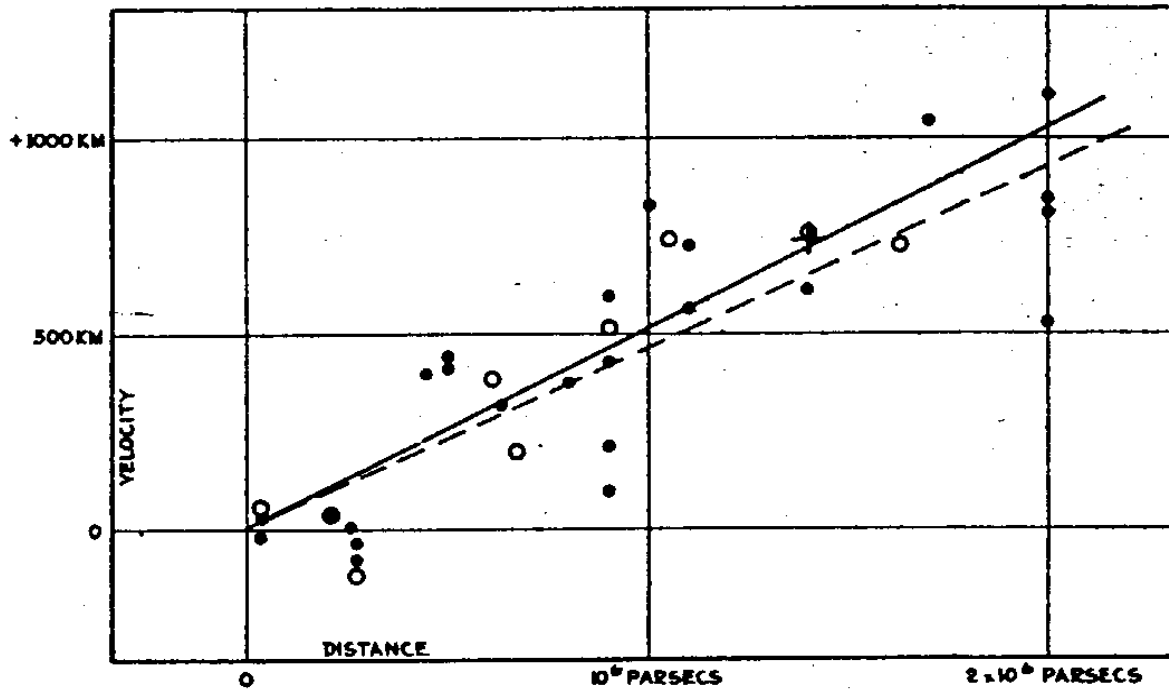
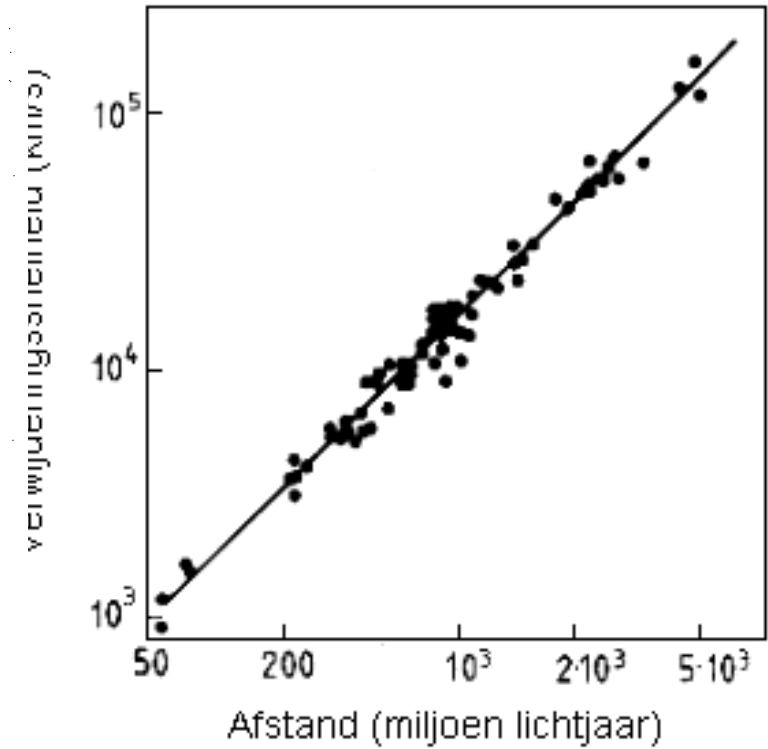


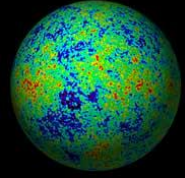
FIGURE 1



De wet van Hubble  $v = H_0 d$

Doppler effect:  $\frac{\Delta\lambda}{\lambda} = \frac{v}{c}$





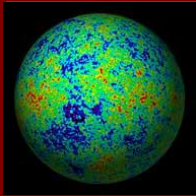
# Het expanderende Heelal

Conclusie: Het Heelal expandeert!

Consequenties:

- Afstanden worden groter!
- Vroeger was de schaal kleiner!
- Het heelal koelt af!
- Vroeger was het kleiner!

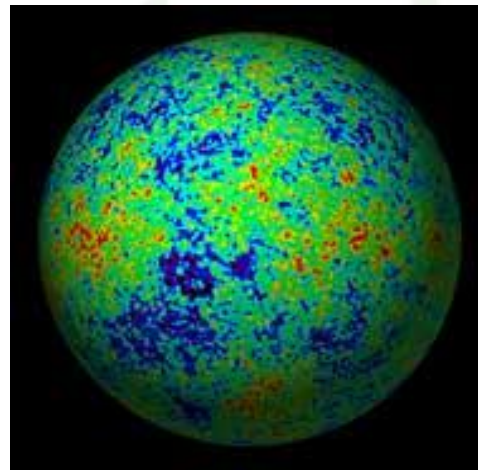
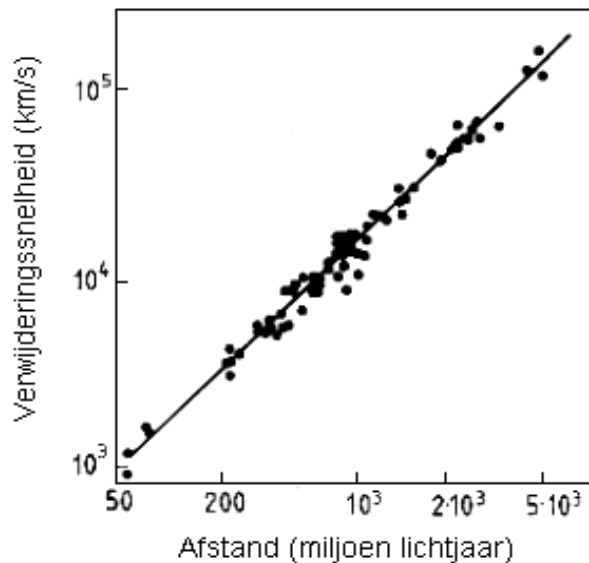
*Het heelal is begonnen in een hete Oerknal!*



# De hete Oerknal

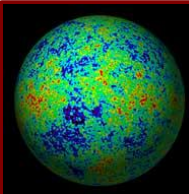
Hete Oerknal model spectaculair bevestigd:

- De expansie van Hubble
- De microgolfachtergrond straling
- Nucleosynthese van de elementen



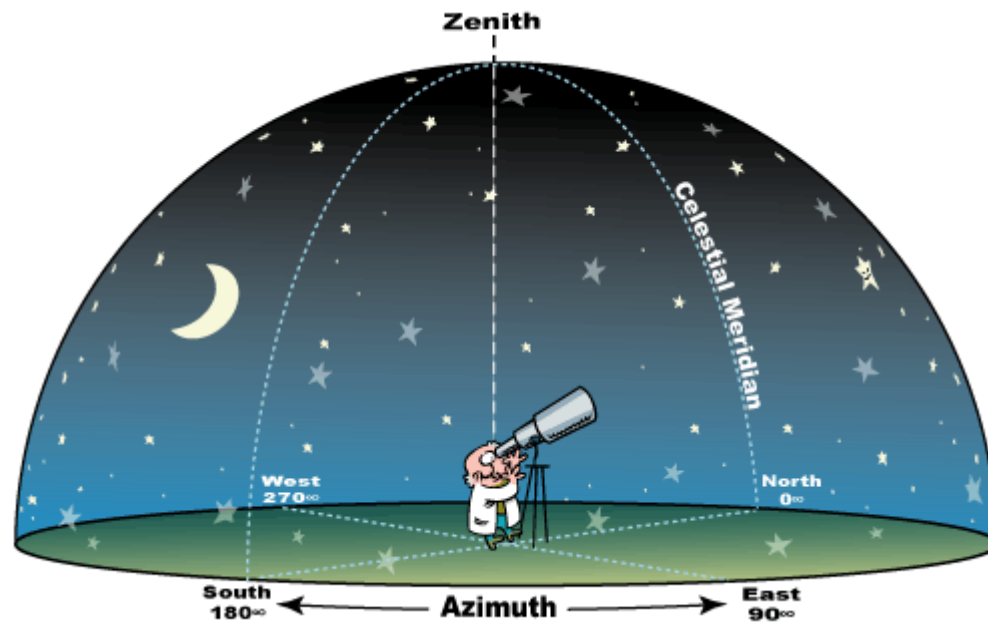
The Astronomer's Periodic Table  
(Ben McCall)

The Astronomer's Periodic Table (Ben McCall)			
H		He	
C	N	O	Ne
Mg	Si	S	Ar
Fe			

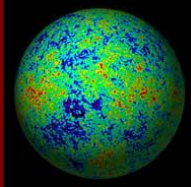


# De Afstandsladder

Kosmologie hangt helemaal samen met afstanden meten in het Heelal





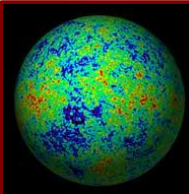


# De Afstandsladder

Aaneenschakeling van methodes:

- Radar ranging
- Parallax
- Variabele sterren (RR Lyrae & Cepheiden)
- Maser afstanden
- Supernovae
- Tully-Fisher & Fundamental Plane
- Hubble expansie

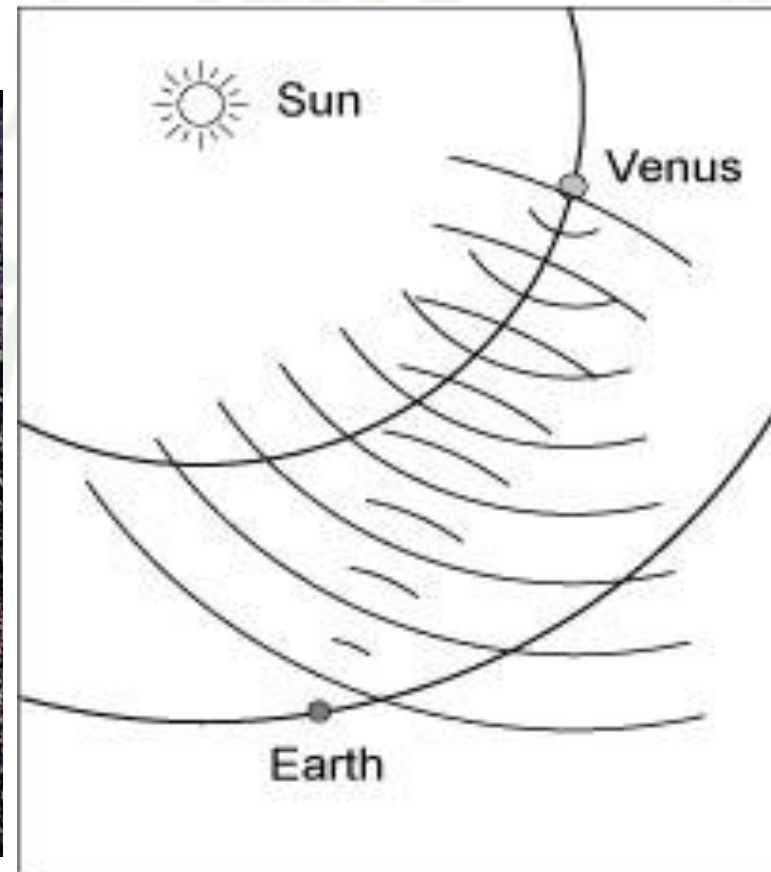
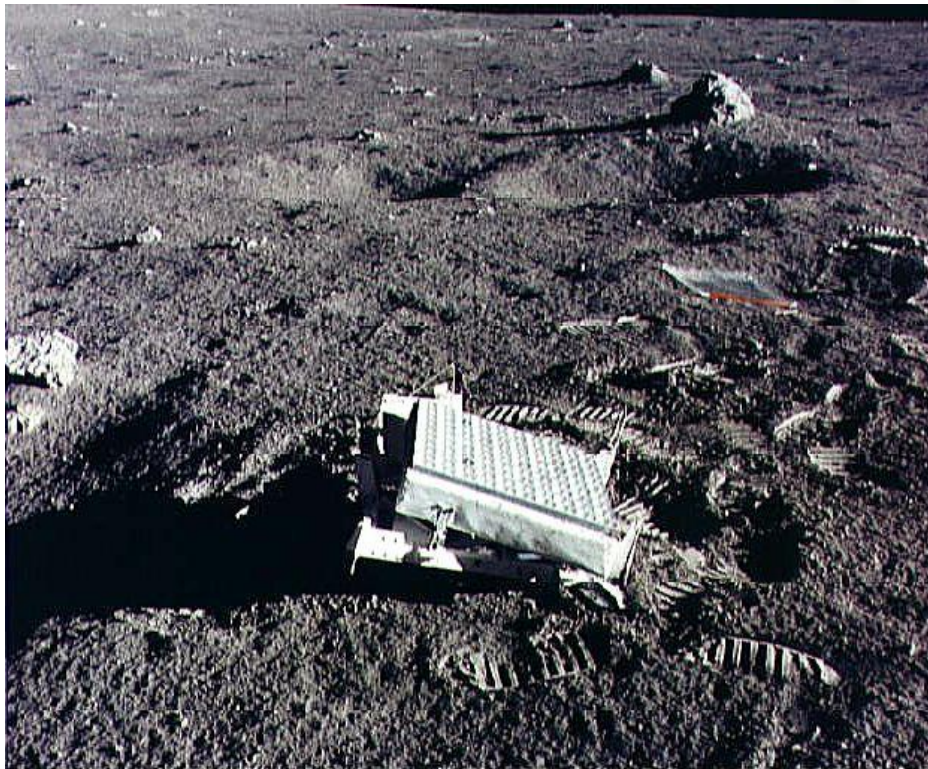




# Radar/Laser Ranging

Uitgangspunt: *lichtsnelheid is constant*

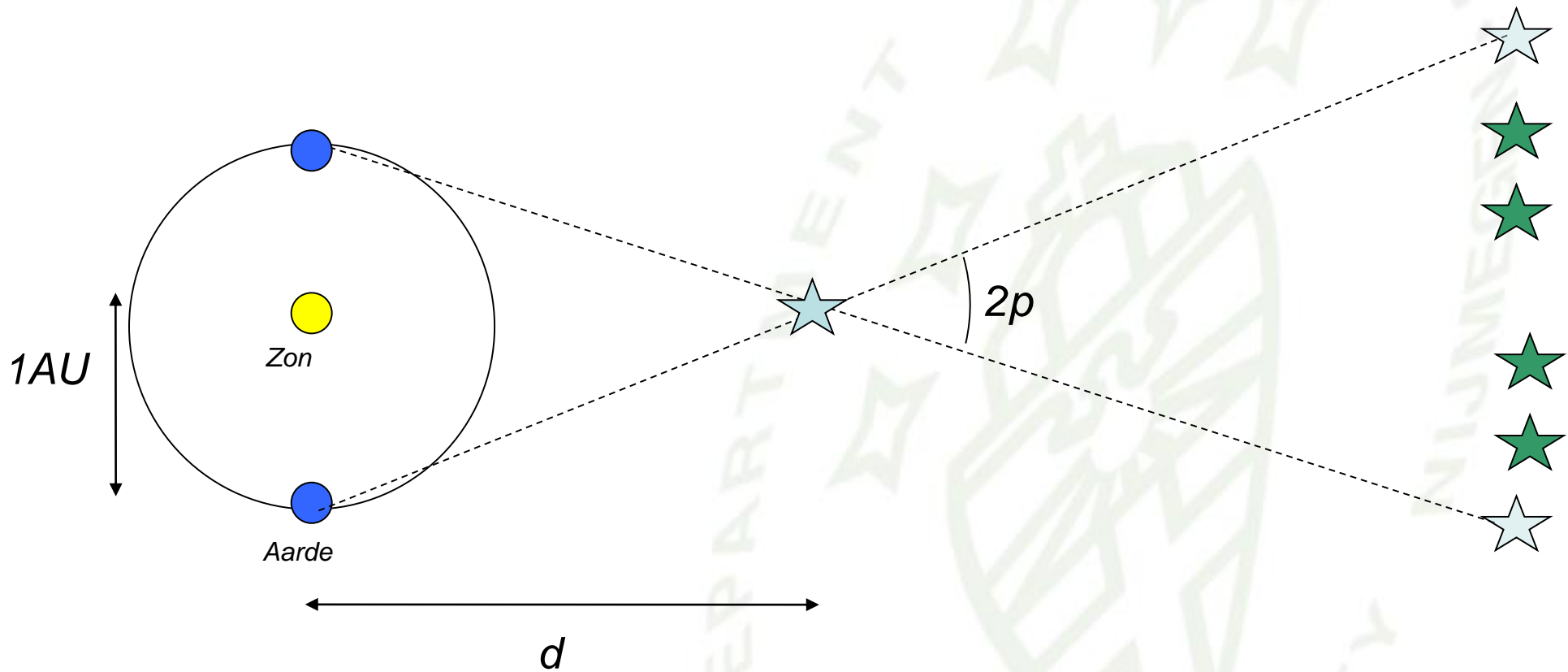
Bereik: Zonnestelsel





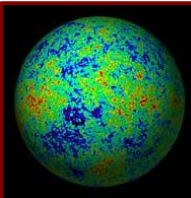


# De parallax meting



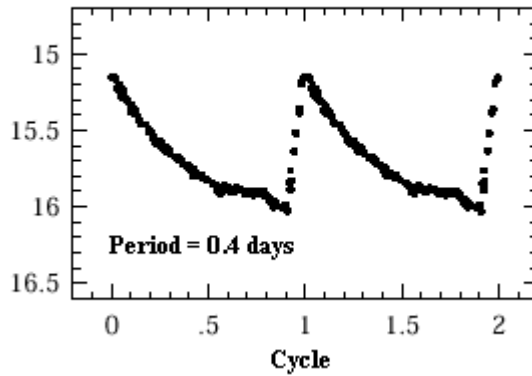
De afstand van de ster:  $\tan p = d / 1 \text{ AU} \approx p$

Dit is de definitie van de afstandsmaat 'parsec': een ster staat op 1 parsec afstand als hij een parallax vertoont van 1 boogseconde.

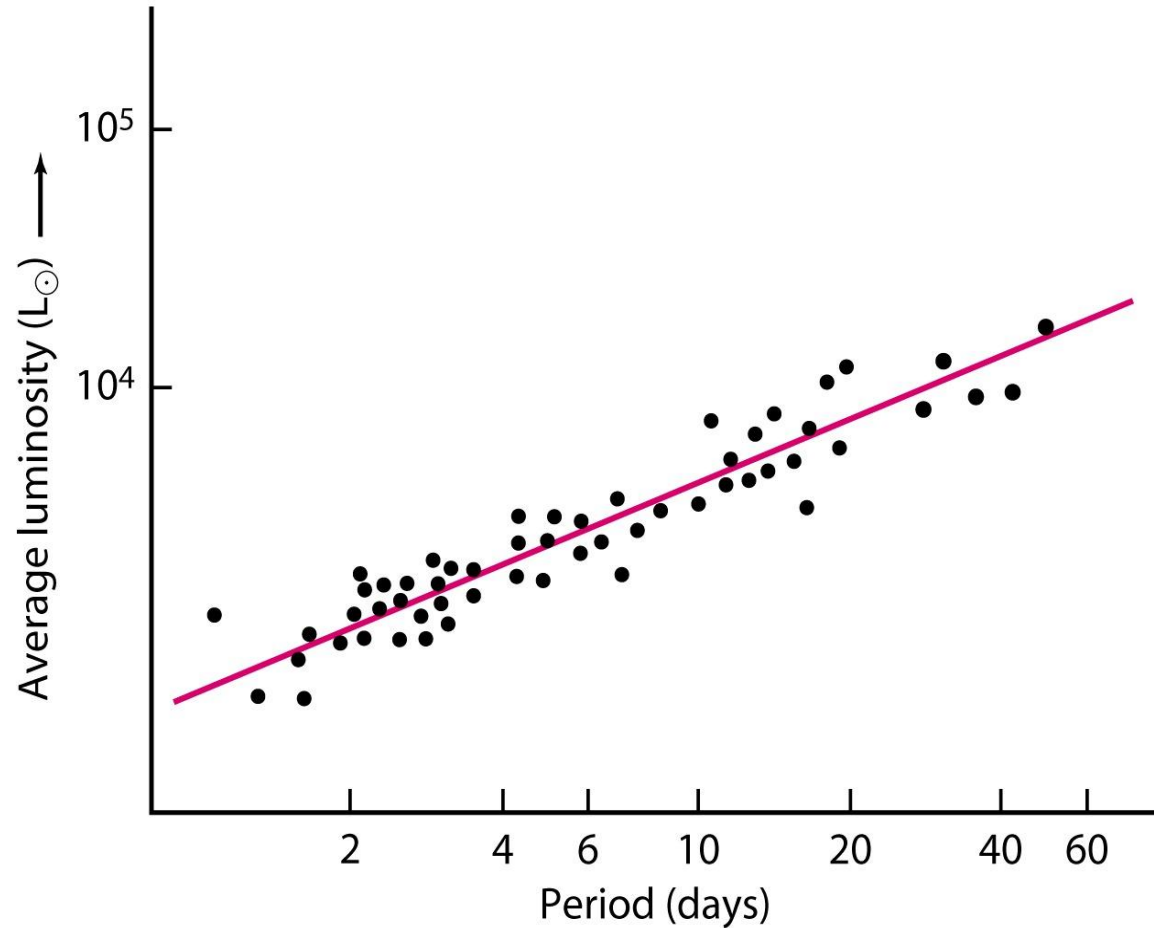
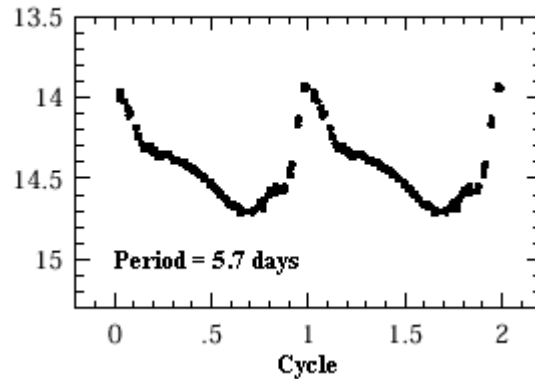


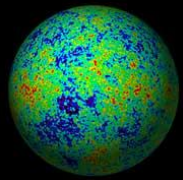
# Variabele Sterren

RR Lyrae Light Variation

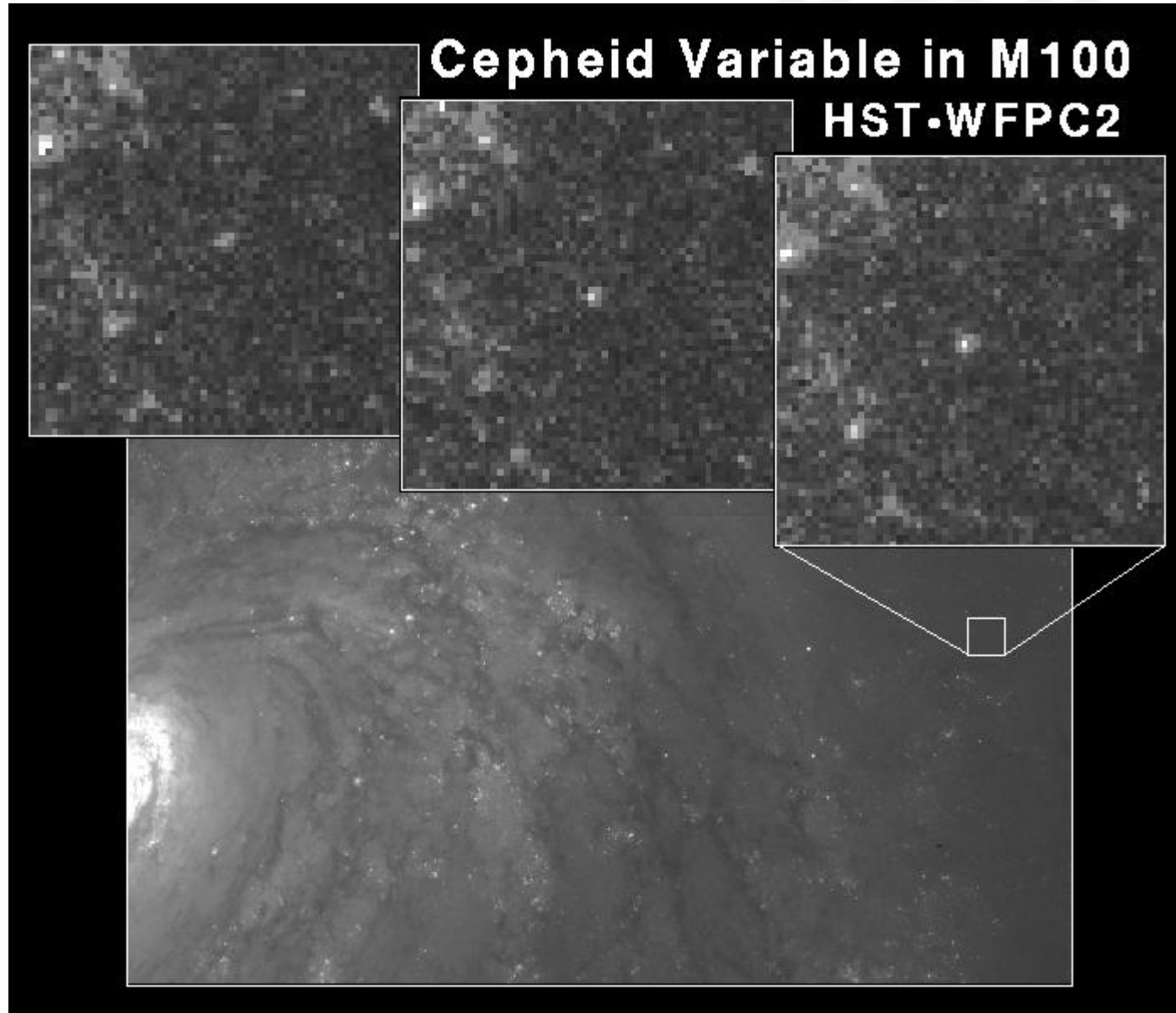


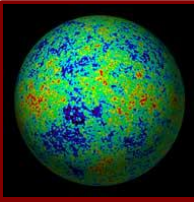
Cepheid Light Variation





# Variabele Sterren





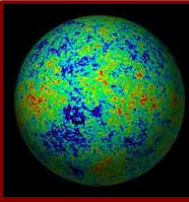
# Supernovae Type Ia

Ontbranding van  
koolstof in witte dwerg  
onder gedegenereerde  
omstandigheden

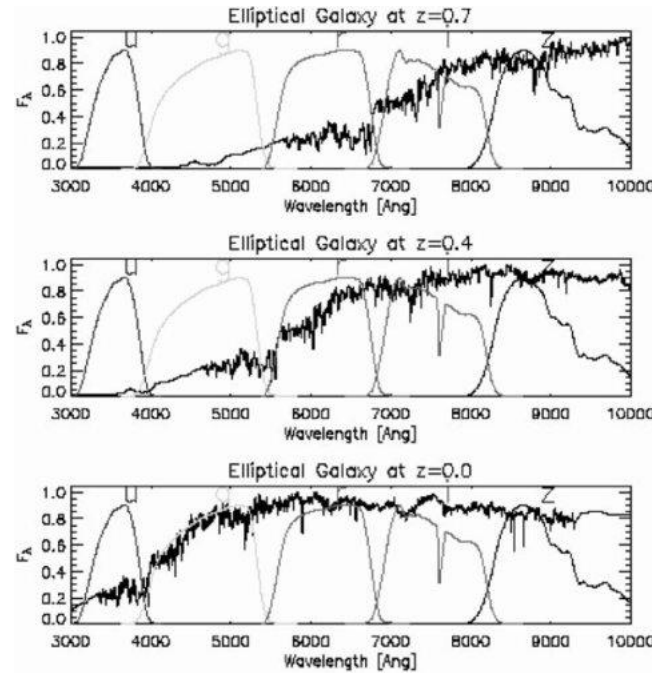
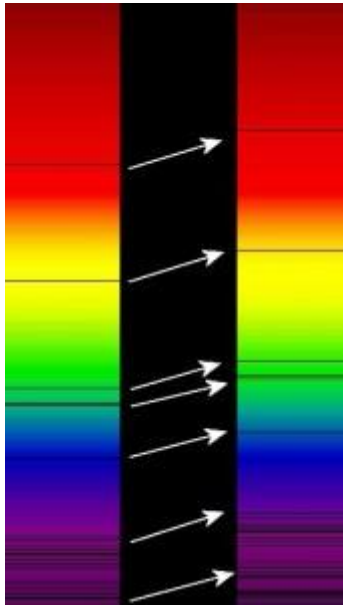
Relatie tussen  
helderheidsverloop en  
helderheid!

→ afstand





# Wet van Hubble



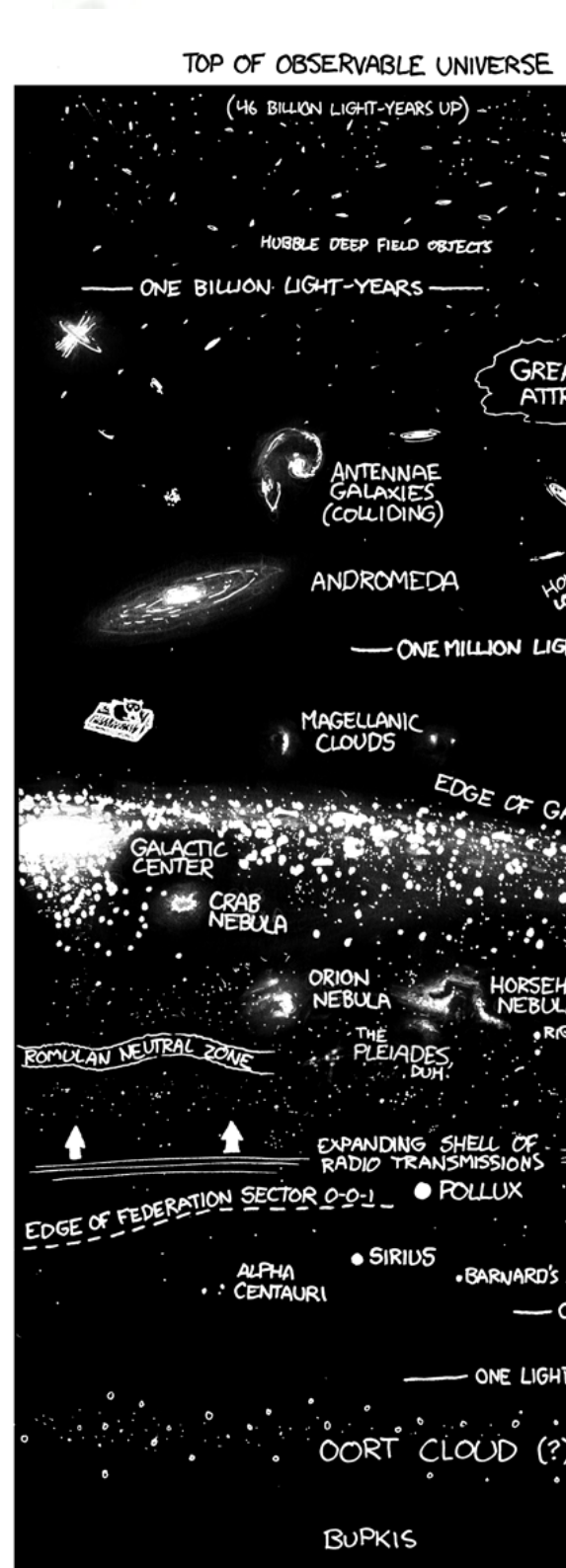
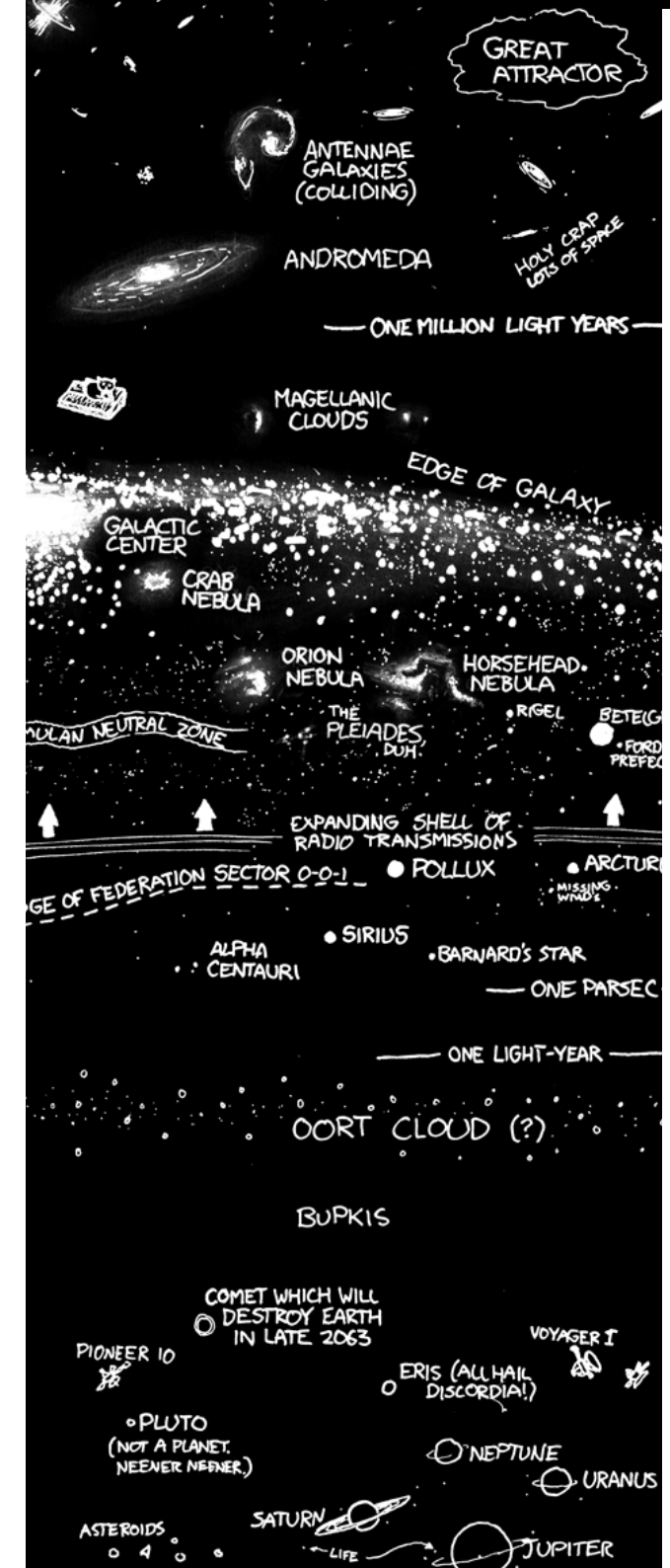
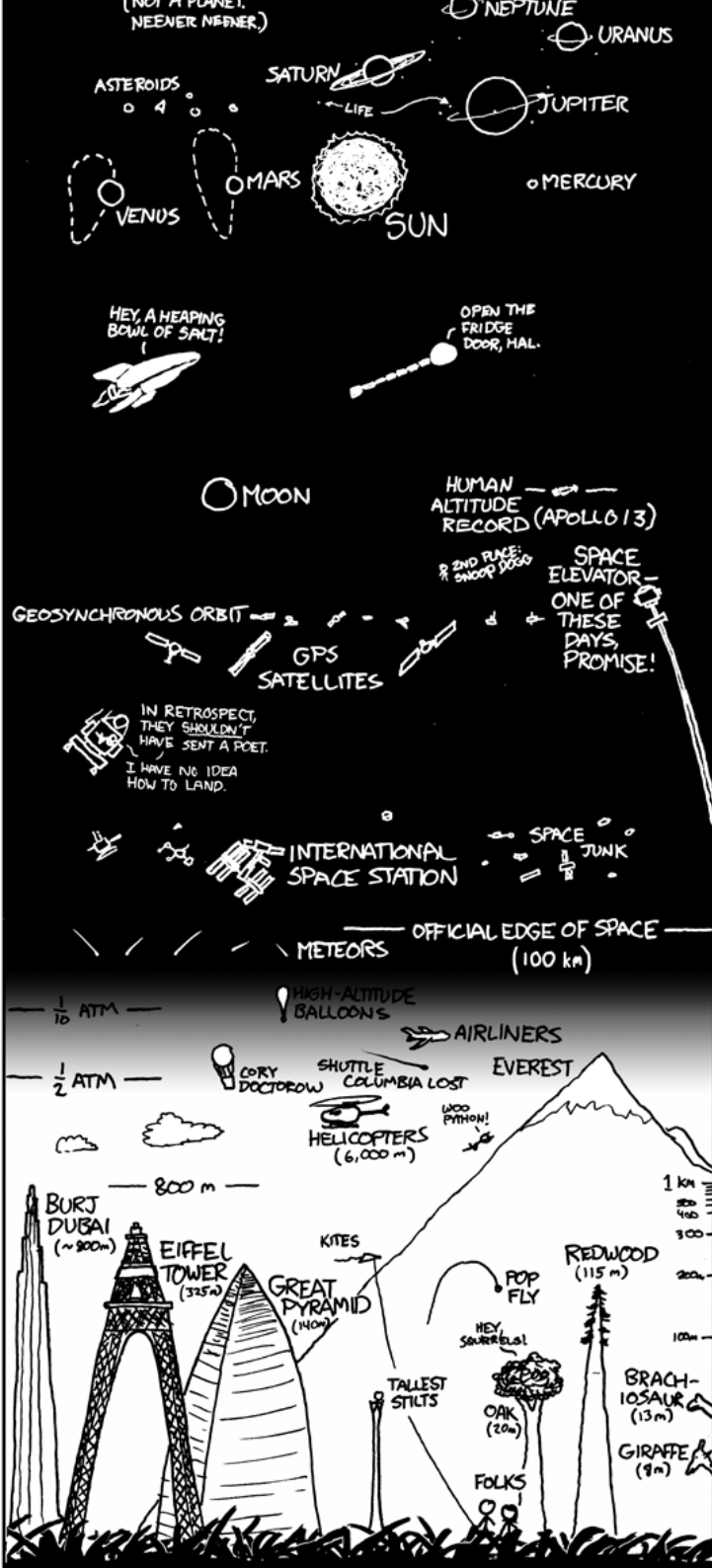
$$z = 1 + \frac{\Delta\lambda}{\lambda_0} = \frac{\lambda - \lambda_0}{\lambda_0}$$

Roodverschuiving

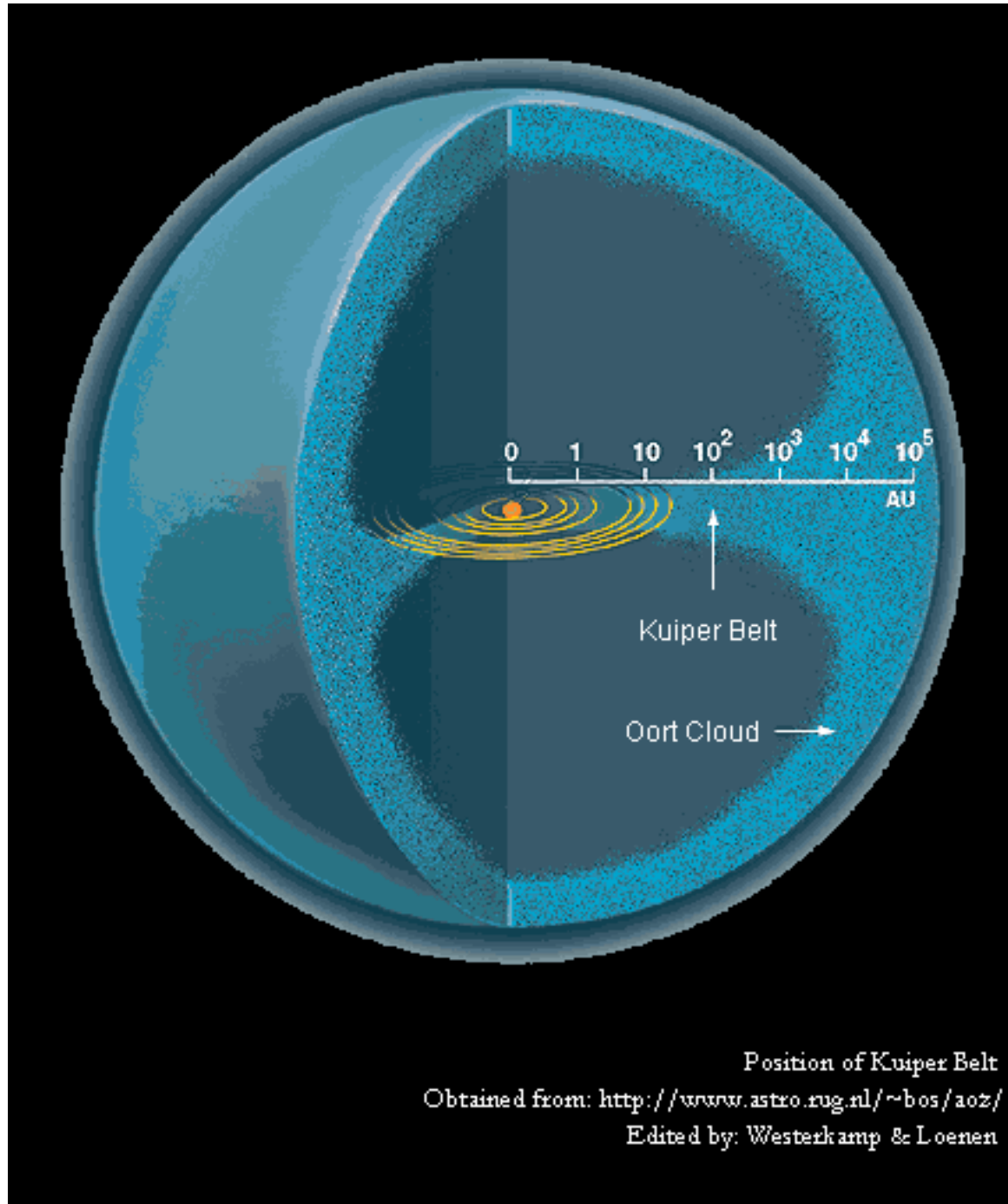
$z+1$  = 'terug-kijktijd' =  $a(t)/a(t_0)$

Fractionele grootte van het Heelal, i.e. op  $z = 5$ : grootte heelal  $1/6$  van nu, etc.





1 lj



100 lj

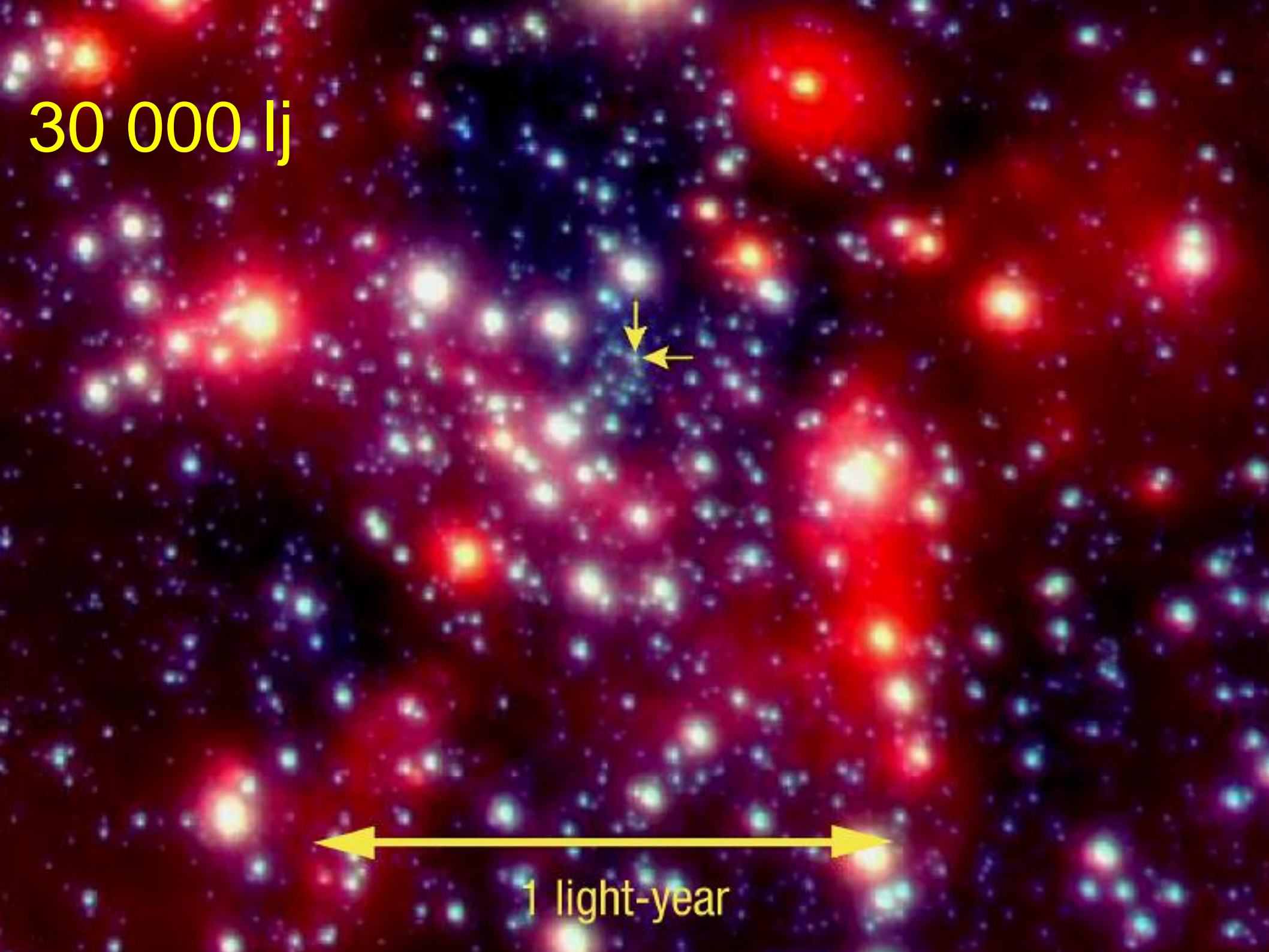




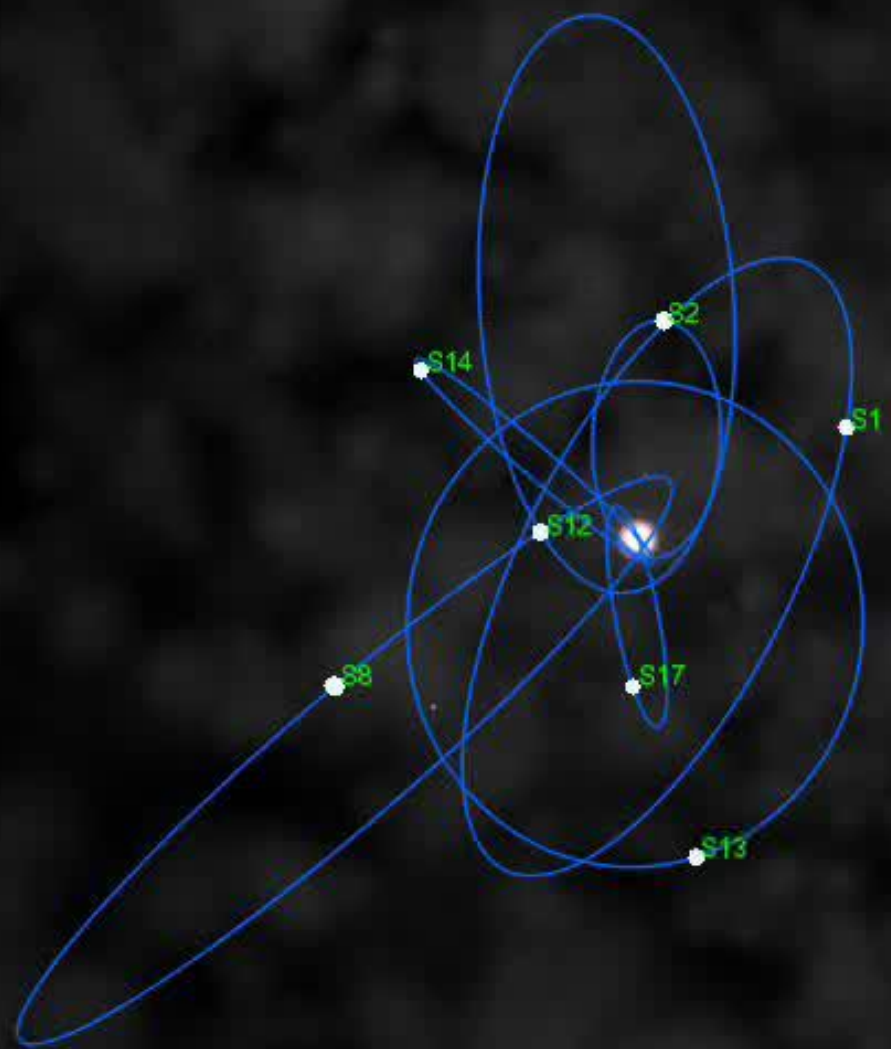
30 000 lj



1 light-year



1993 09 09 13:58:59 UTC  
45000000\* faster



**| -10 light days - |**

Speed: 0.000 m/s

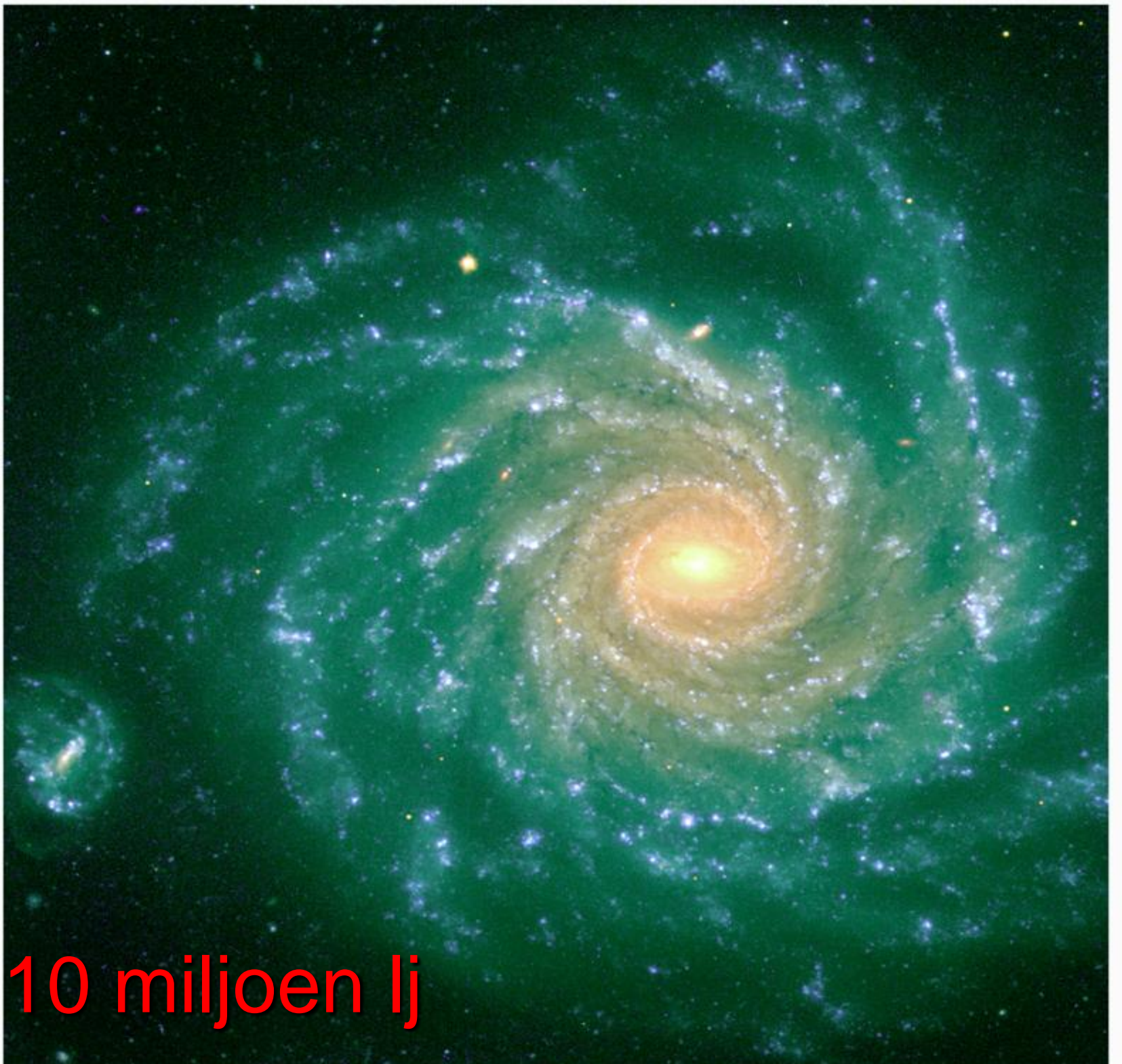
Follow GC  
FOV: 13° 59' 60.0" (1.00x)



# Andromeda, ruim 2 miljoen lj







10 miljoen lj

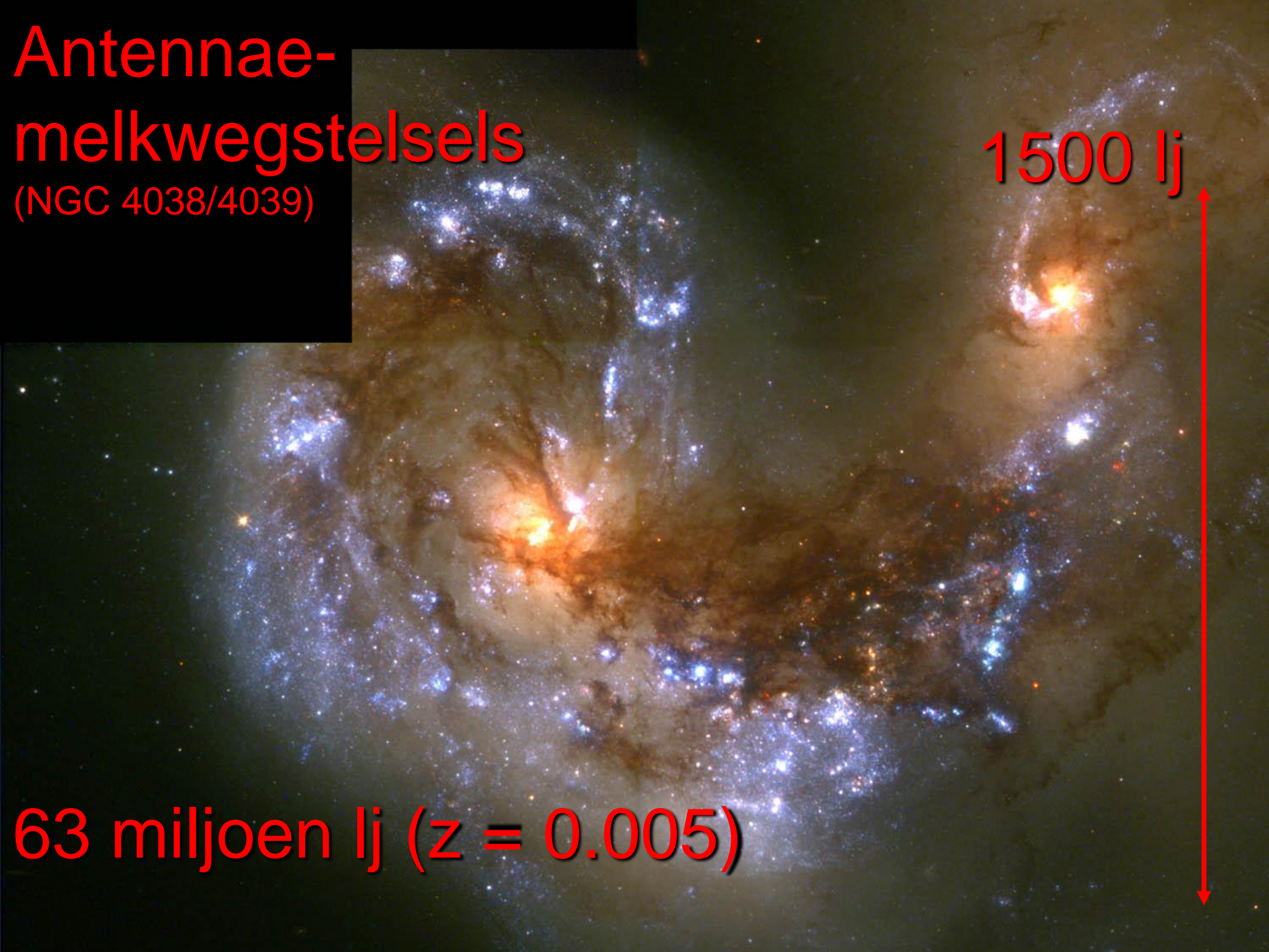


# Antennae- melkwegstelsels (NGC 4038/4039)

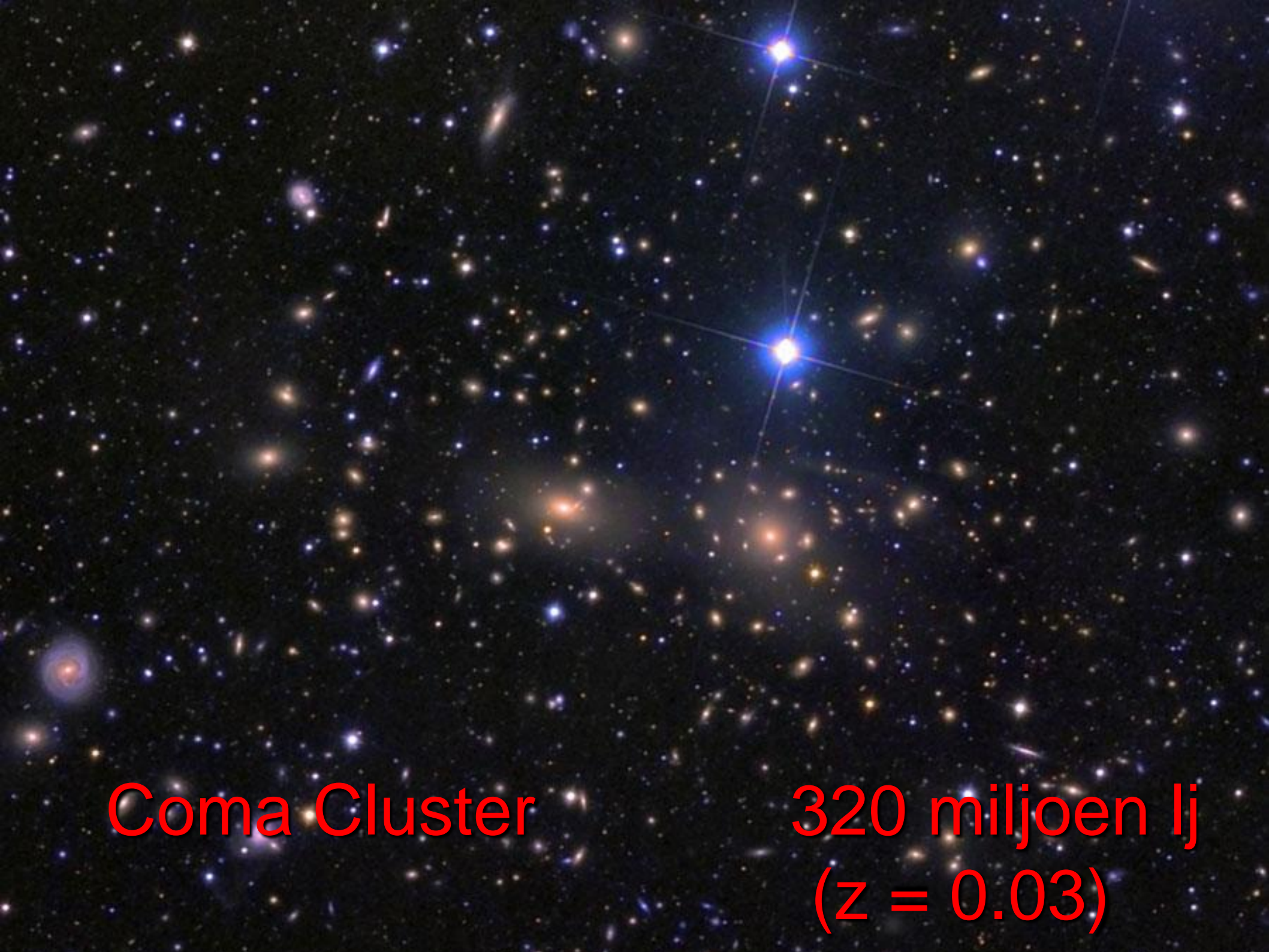
1500 lj



63 miljoen lj ( $z = 0.005$ )



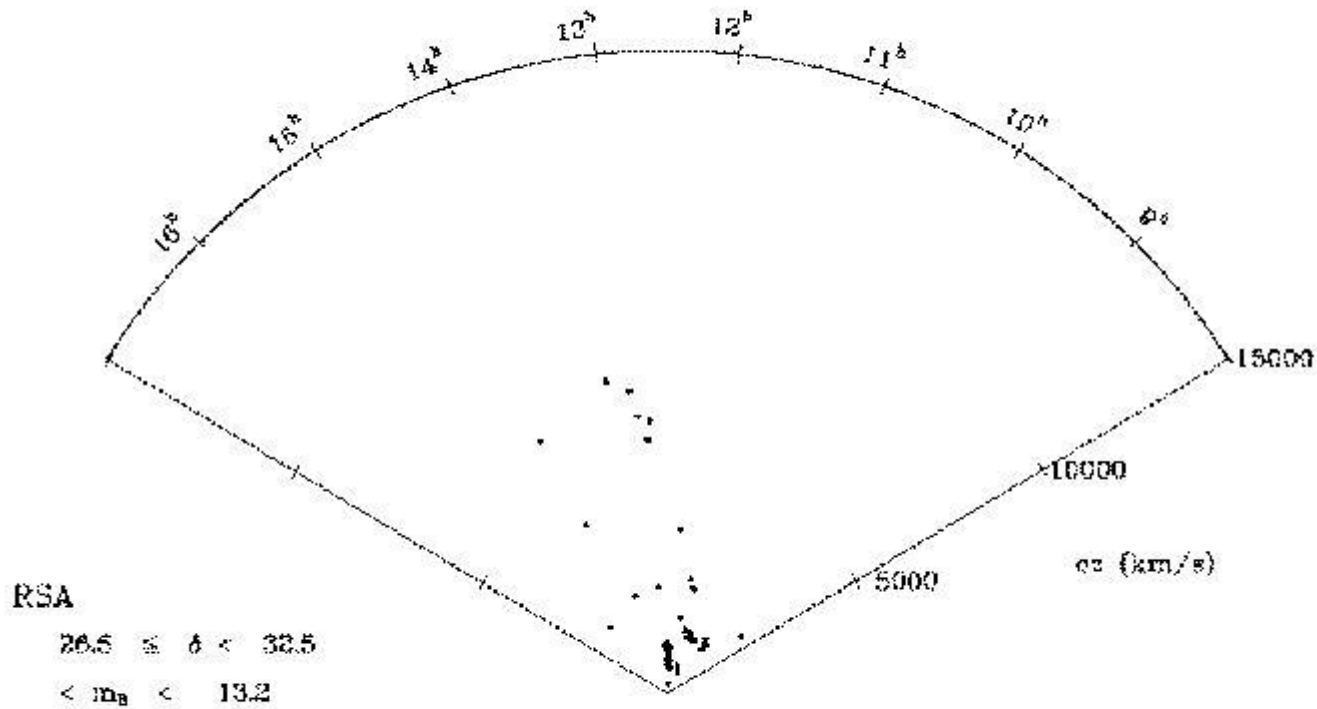




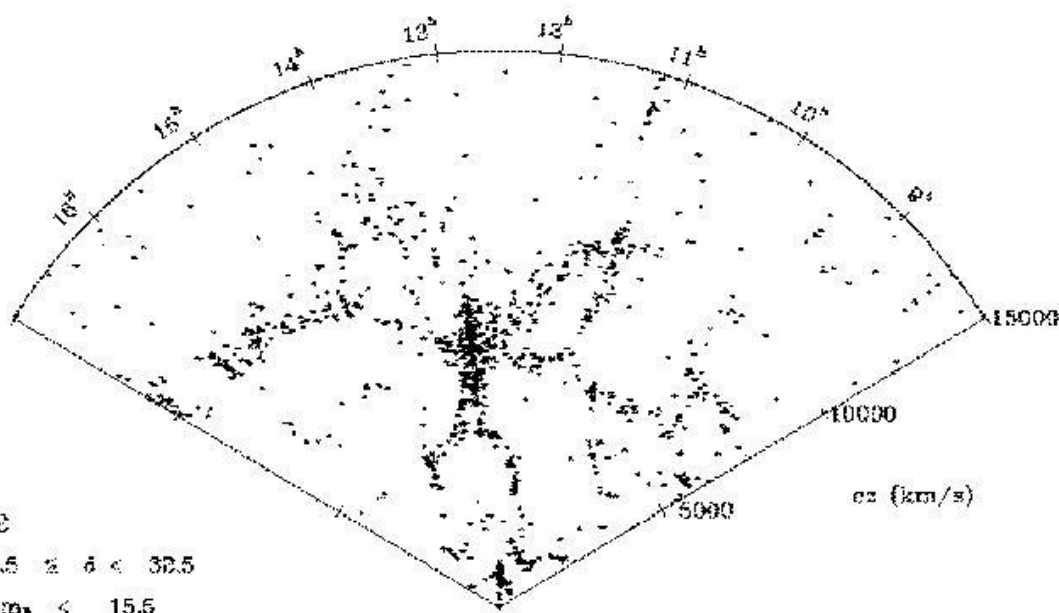
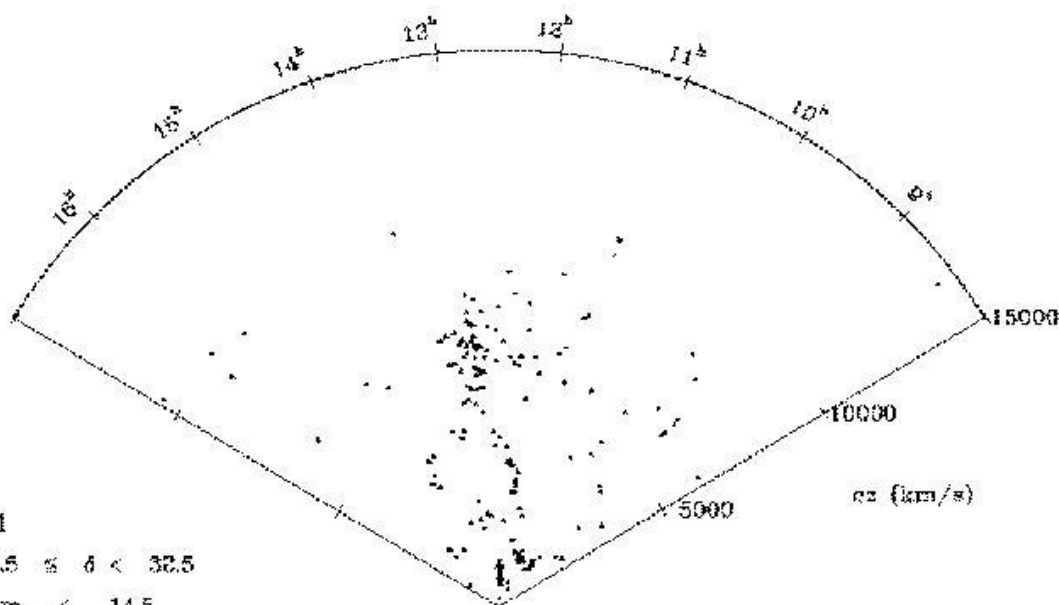
Coma Cluster

320 miljoen lj  
( $z = 0.03$ )

# Jaren '70: surveys





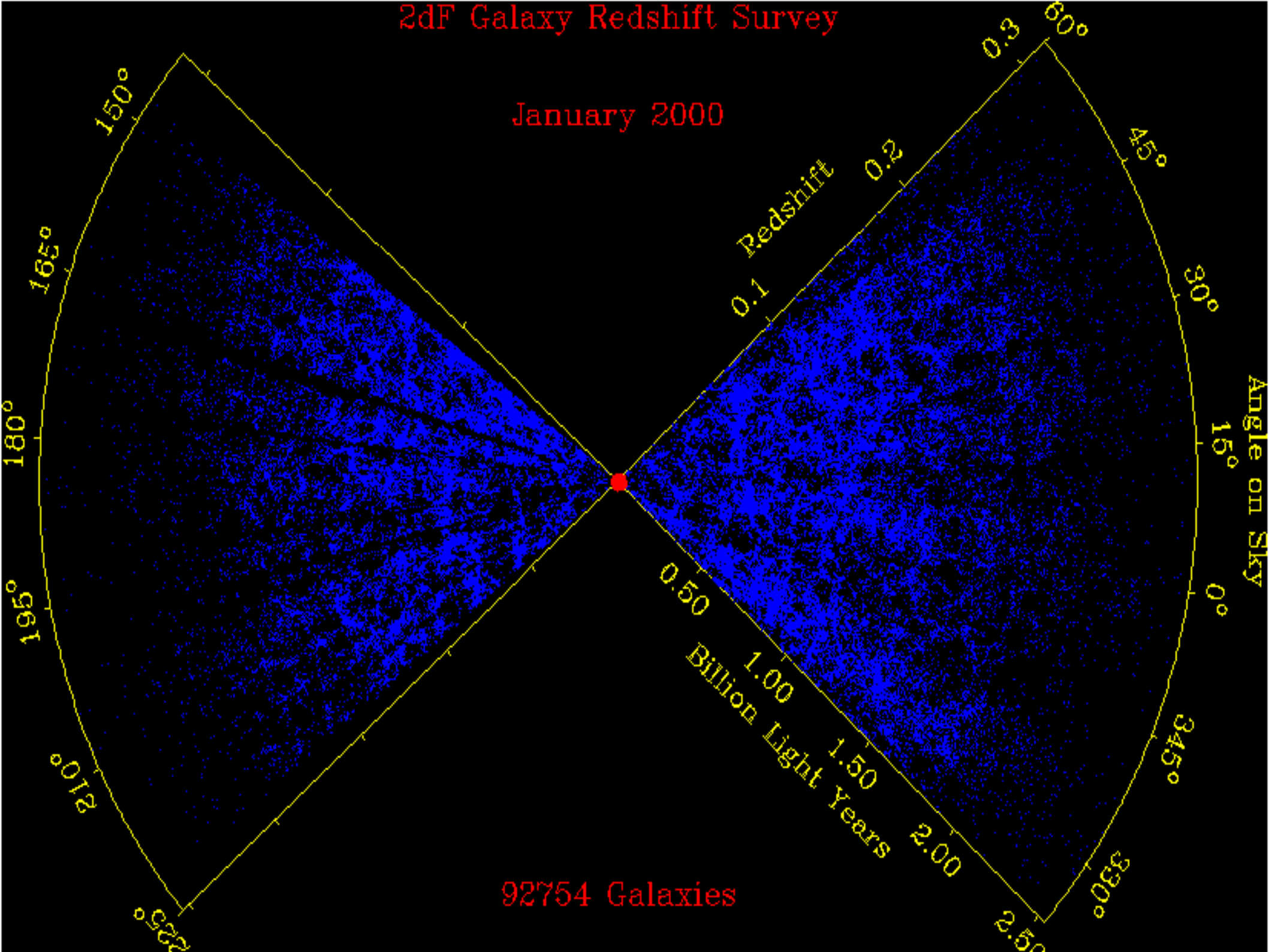


SCHAUMMEGEN · 53

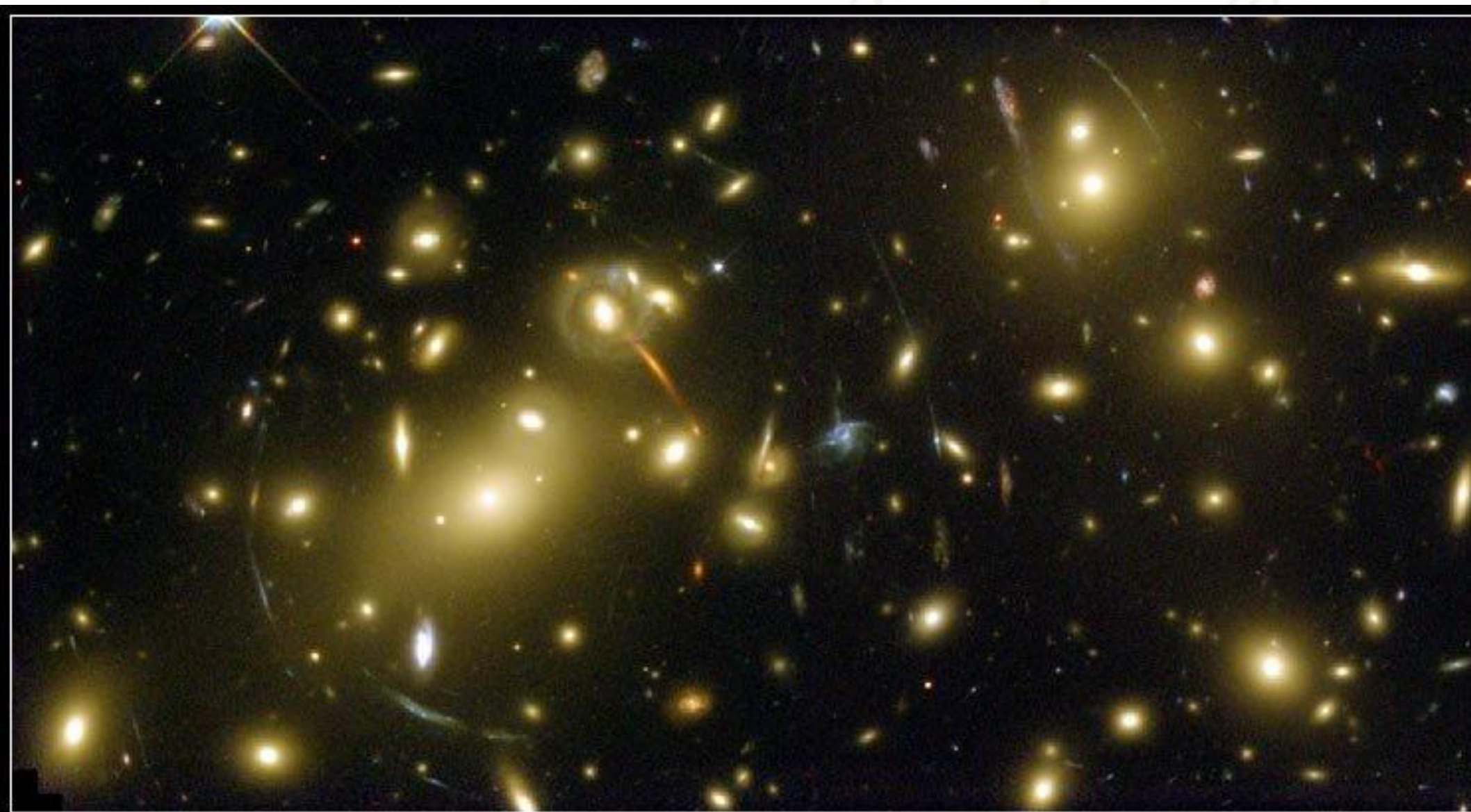
Schuimachtige structuur!

# 2dF Galaxy Redshift Survey

January 2000



$Z = 0.17$



**Galaxy Cluster Abell 2218**

**HST • WF**

NASA, A. Fruchter and the ERO Team (STScI) • STScI-PRC00-08



$z = 0.7$

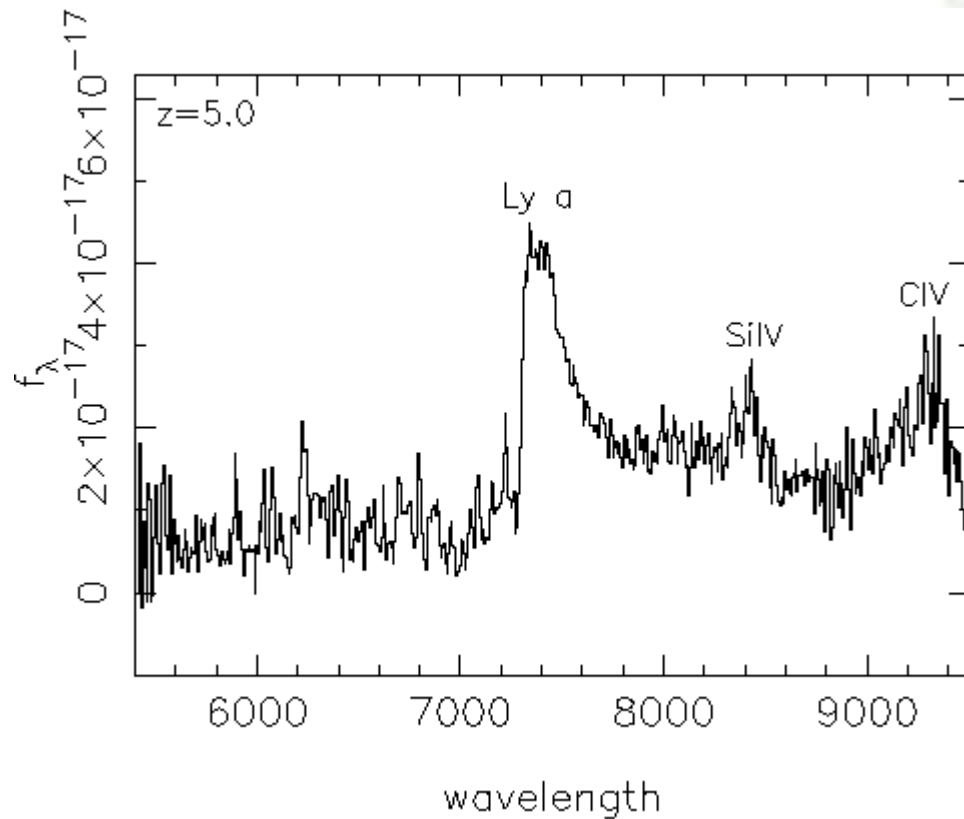




# Sloan Digital Sky Survey

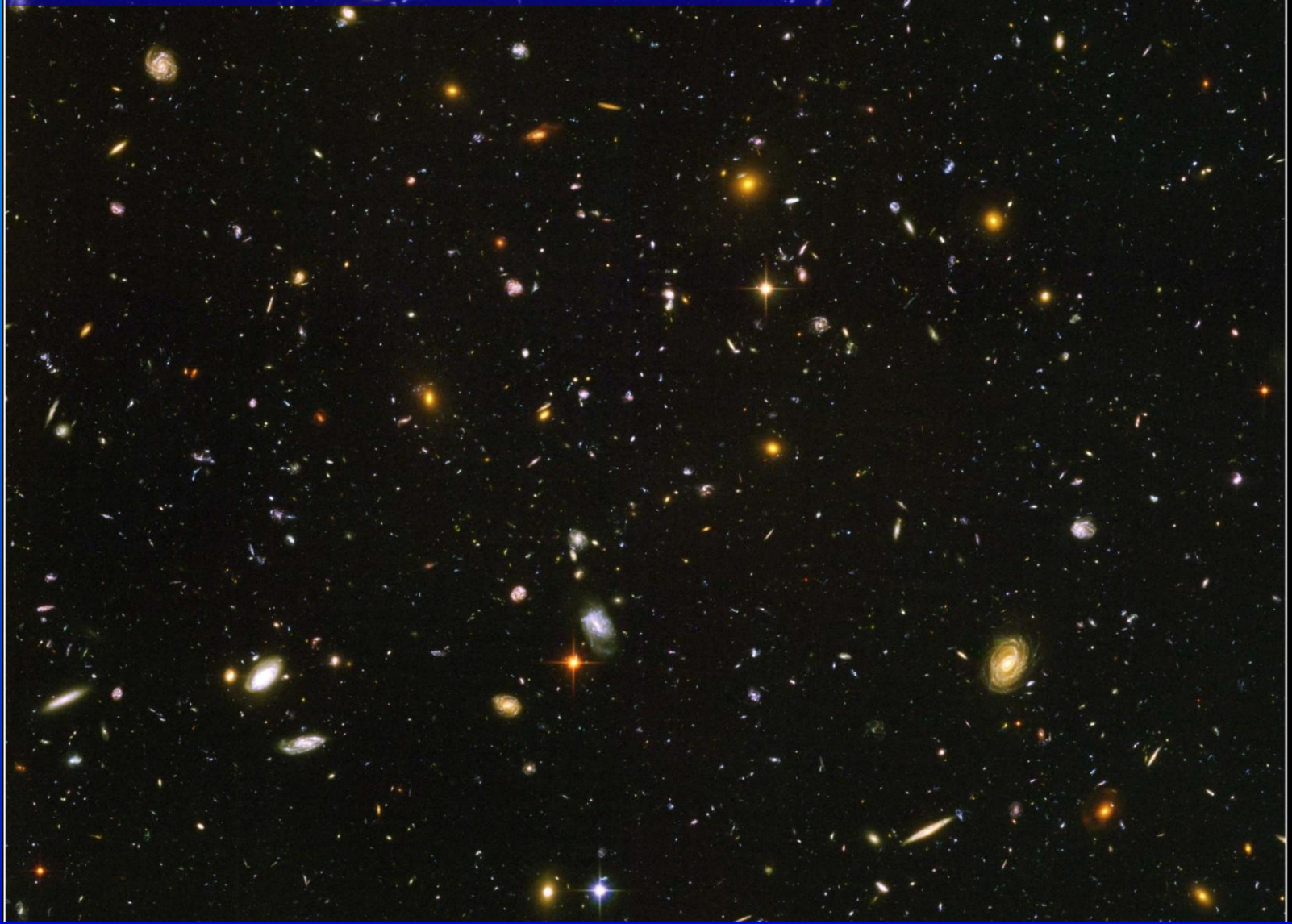
$z = 4.75$  Quasar

$z = 5.0$  Quasar

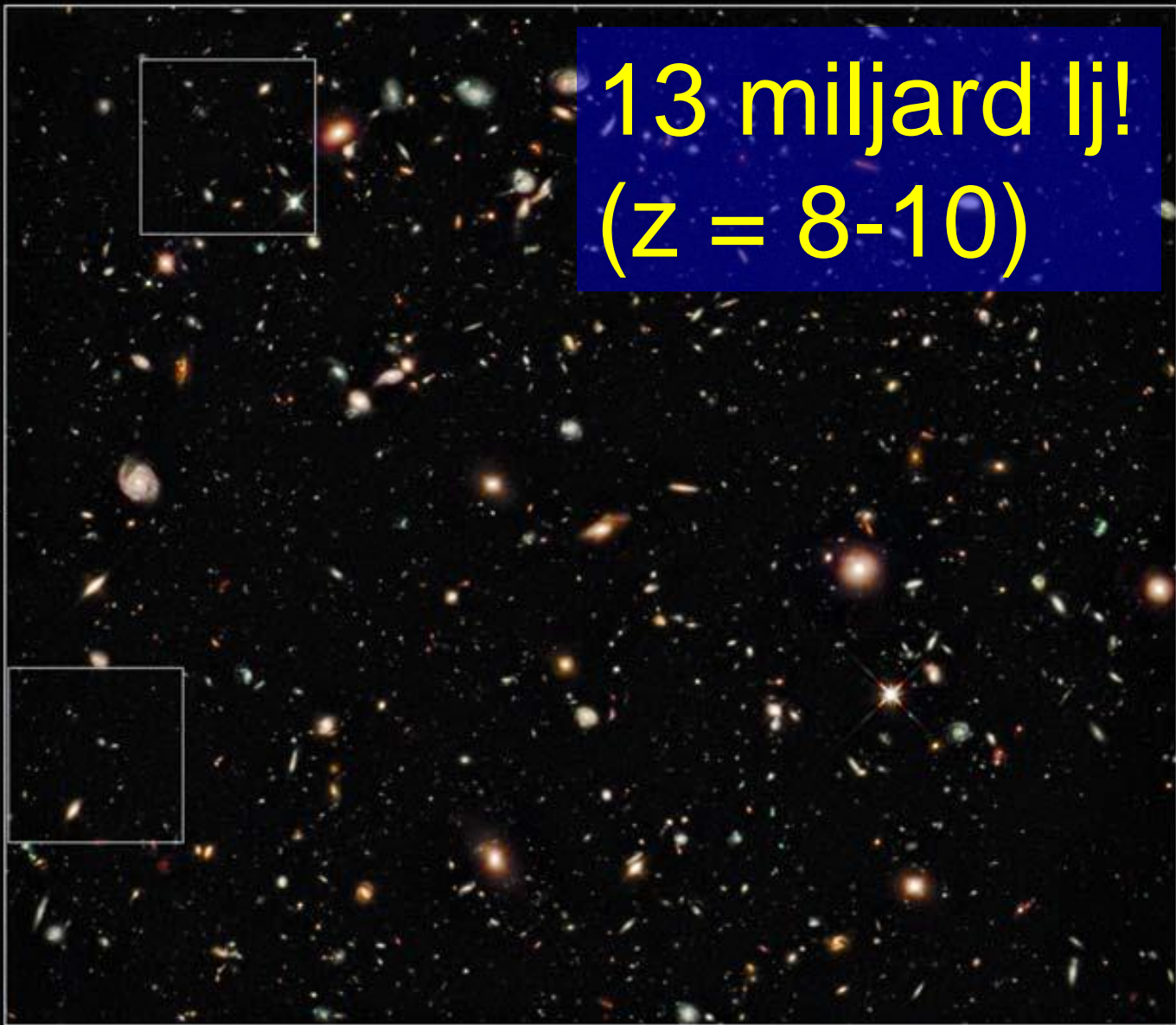
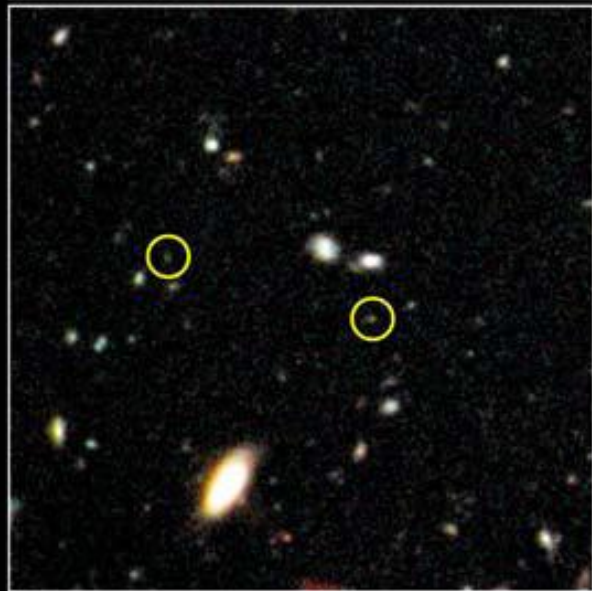
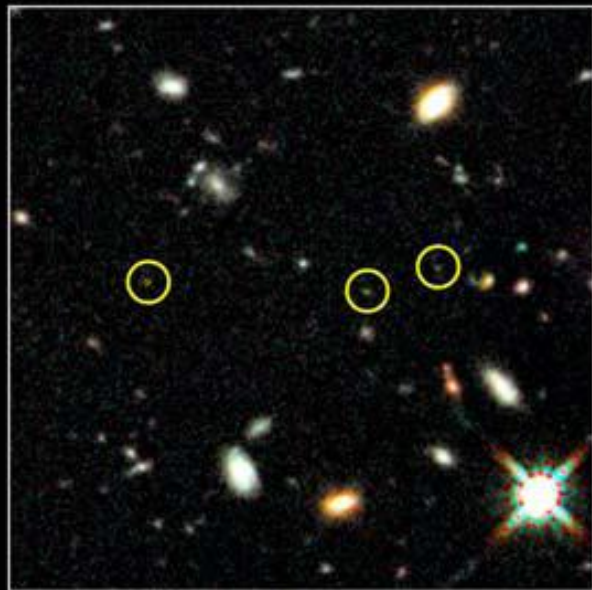




Hubble Deep Field  
10 miljard lj  
( $z = 6$ )

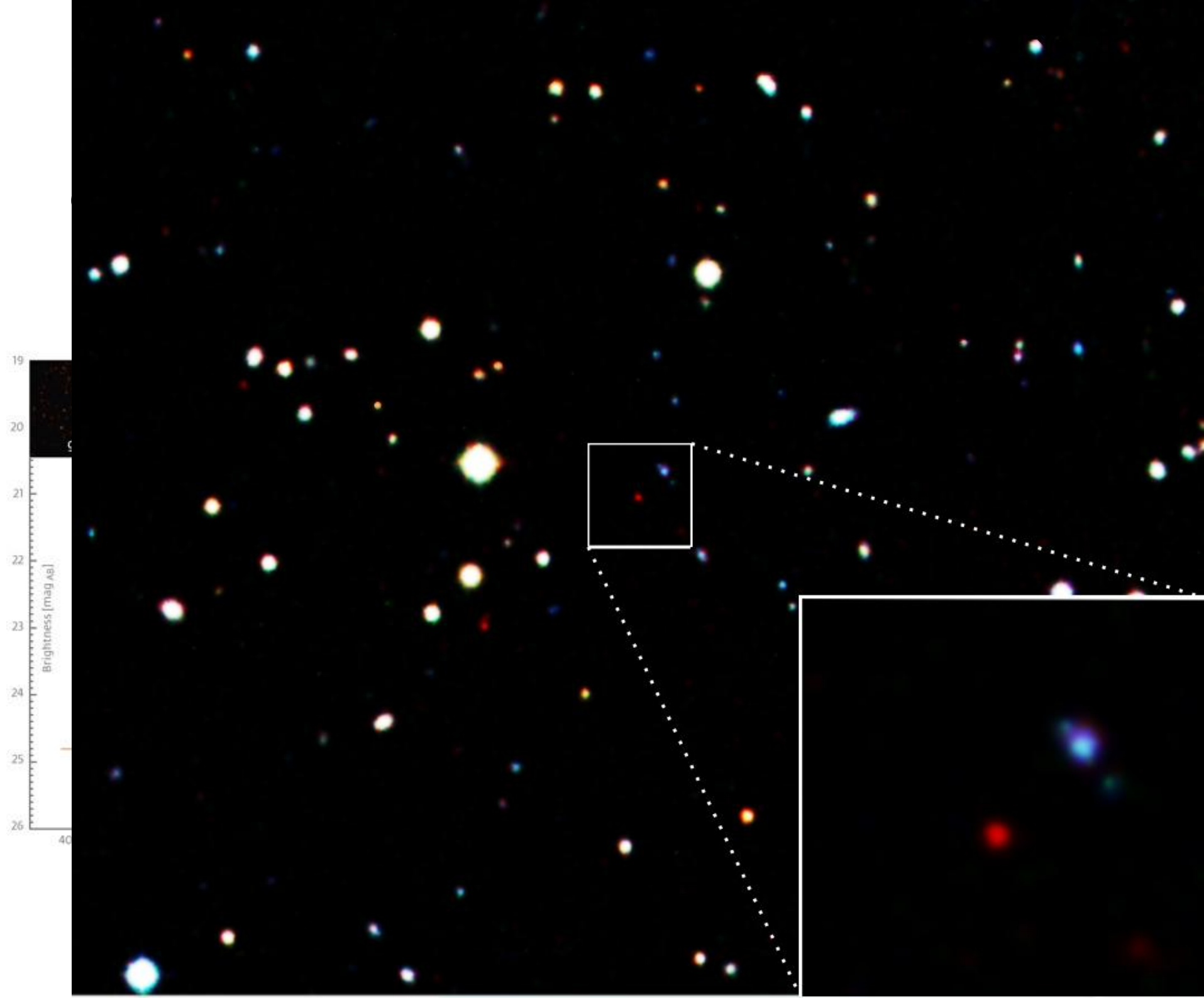






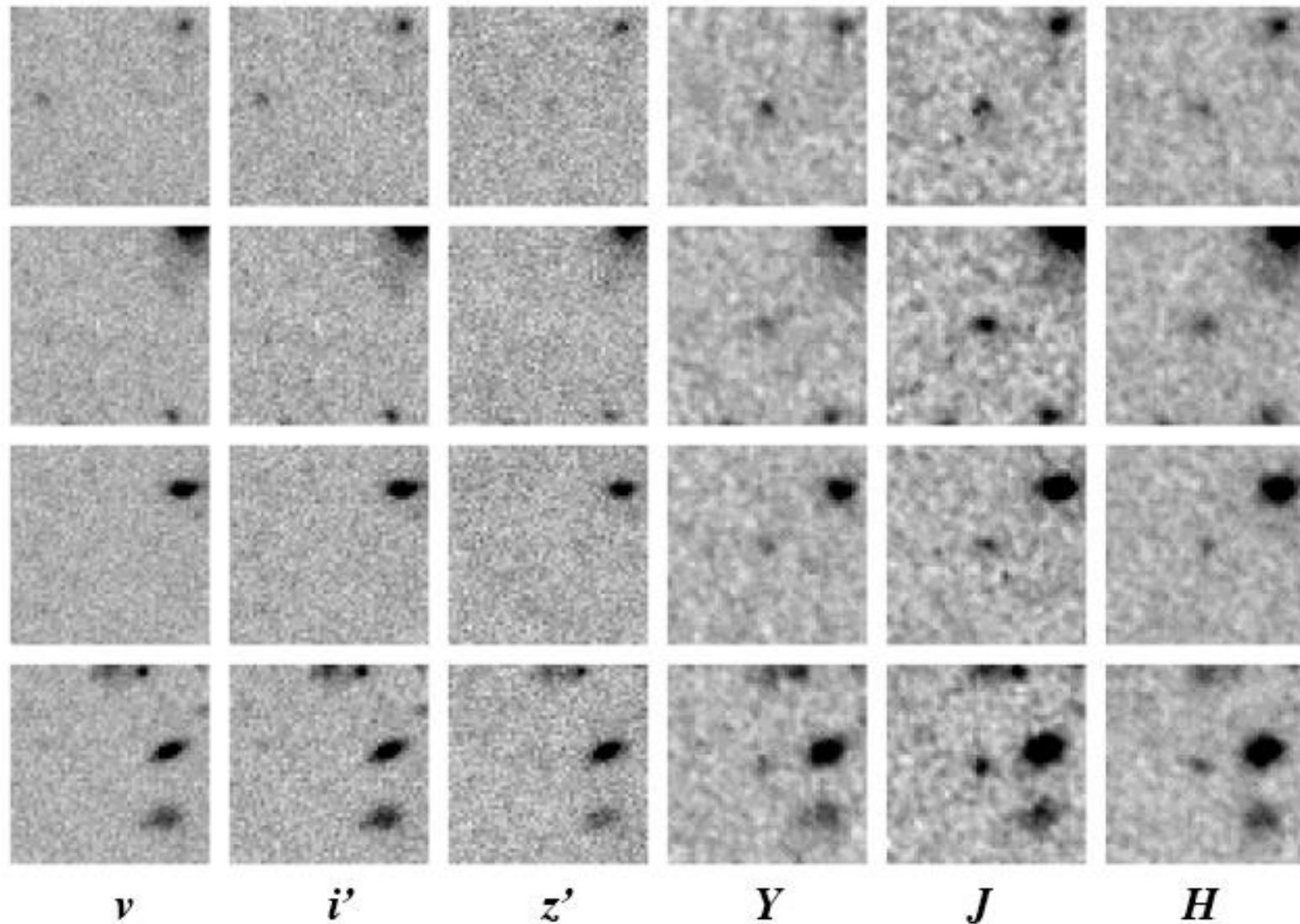
13 miljard lj!  
( $z = 8-10$ )

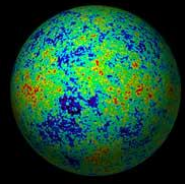
Hubble Ultra Deep Field • Infrared  
*Hubble Space Telescope* • WFC3/IR





# J drop outs





# Hoe ver kunnen we terug?

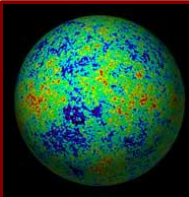
- Terug in tijd: heelal steeds heter
- Gamow, Alpher (1948)
  - Mogelijke verklaring elementen He etc. (temperatuur voor H fusie  $\sim 10^9$  K)  $\rightarrow$  volgende week
  - Bij temperatuur  $\sim 10\ 000$  K  $\rightarrow$  waterstof ioniseert ( $p + e^-$ )
  - Daarvoor: straling “gevangen” door verstrooiing met electronen
  - Daarna: straling beweegt vrij door Heelal
  - Door uitzetting Heelal worden fotonen “langer”. Temperatuur daalt, nu  $T \sim 5$ K
- Nog zichtbaar?
  - Jaren 60: Dicke, Peebles (Princeton) gaan op zoek



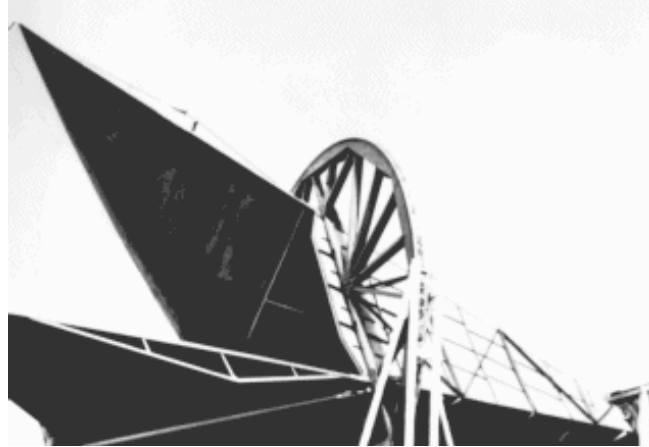
Gamow



Alpher



# 1964 Penzias & Wilson



No. 1, 1965

LETTERS TO THE EDITOR

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high pressure, such as the zero-mass scalar, capable of speeding the universe through the period of helium formation. To have a closed space, an energy density of  $2 \times 10^{-29}$  gm/cm<sup>3</sup> is needed. Without a zero-mass scalar, or some other "hard" interaction, the energy could not be in the form of ordinary matter and may be presumed to be gravitational radiation (Wheeler 1958).

One other possibility for closing the universe, with matter providing the energy content of the universe, is the assumption that the universe contains a net electron-type neutrino abundance (in excess of antineutrinos) greatly larger than the nucleon abundance. In this case, if the neutrino abundance were so great that these neutrinos are degenerate, the degeneracy would have forced a negligible equilibrium neutron abundance in the early, highly contracted universe, thus removing the possibility of nuclear reactions leading to helium formation. However, the required ratio of lepton to baryon number must be  $> 10^9$ .

We deeply appreciate the helpfulness of Drs. Penzias and Wilson of the Bell Telephone Laboratories, Crawford Hill, Holmdel, New Jersey, in discussing with us the result of their measurements and in showing us their receiving system. We are also grateful for several helpful suggestions of Professor J. A. Wheeler.

R. H. DICKE

## A MEASUREMENT OF EXCESS ANTENNA TEMPERATURE AT 4080 Mc/s

Measurements of the effective zenith noise temperature of the 20-foot horn-reflector antenna (Crawford, Hogg, and Hunt 1961) at the Crawford Hill Laboratory, Holmdel, New Jersey, at 4080 Mc/s have yielded a value about **3.5° K** higher than expected. This excess temperature is, within the limits of our observations, **isotropic**, unpolarized, and

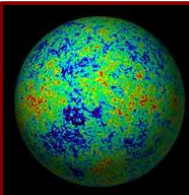


— 1964, in *Relativity, Groups and Topology*, ed C. DeWitt and B. DeWitt (New York: Gordon & Breach).  
Zel'dovich, Ya. B. 1962, *Soviet Phys.—J.E.T.P.*, 14, 1143.

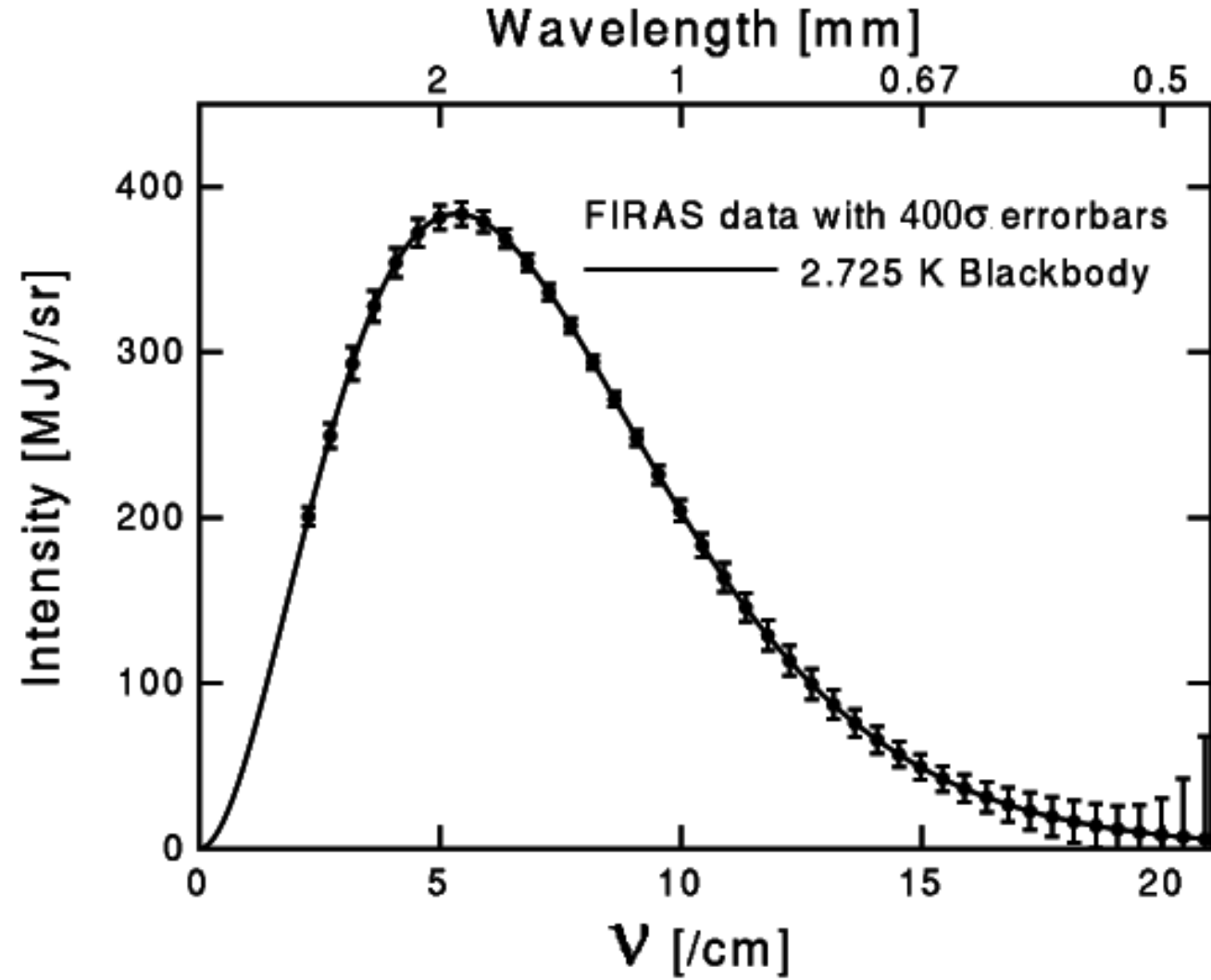
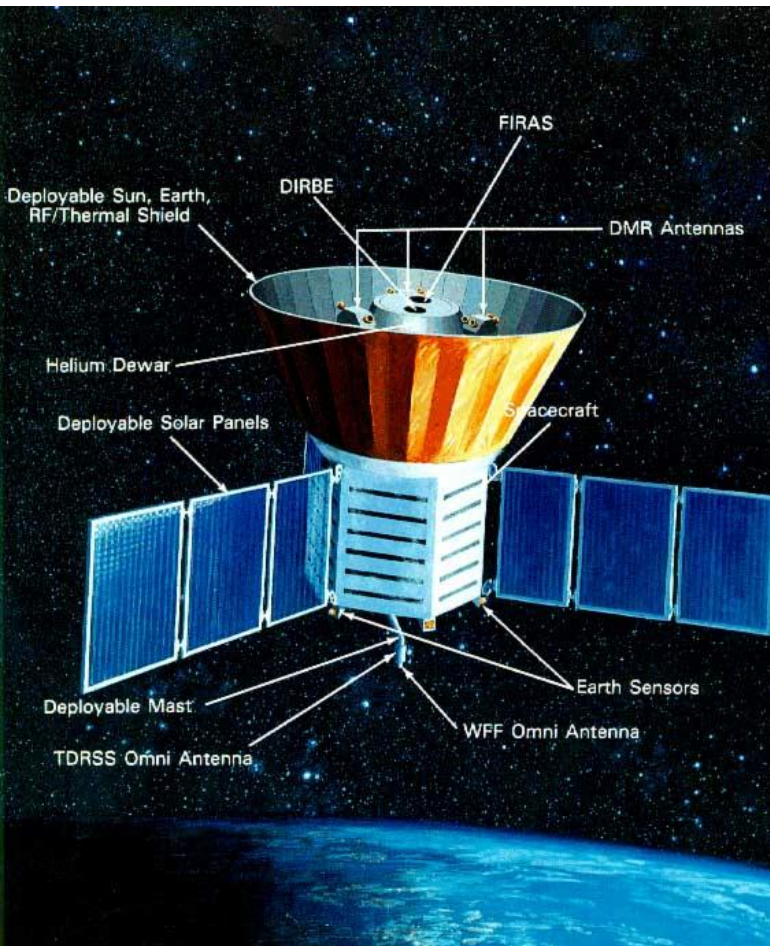
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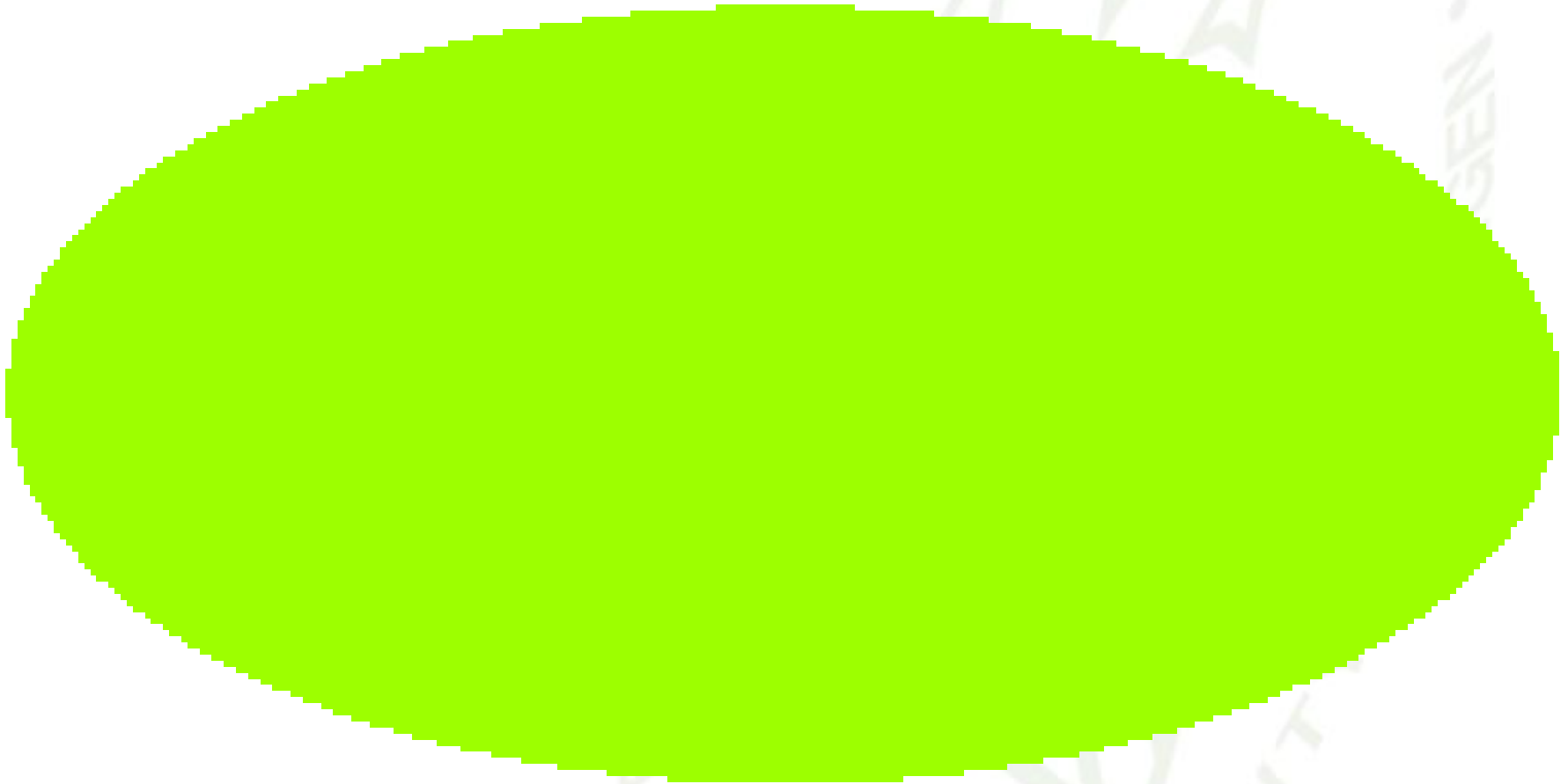


# Jaren '90: COBE



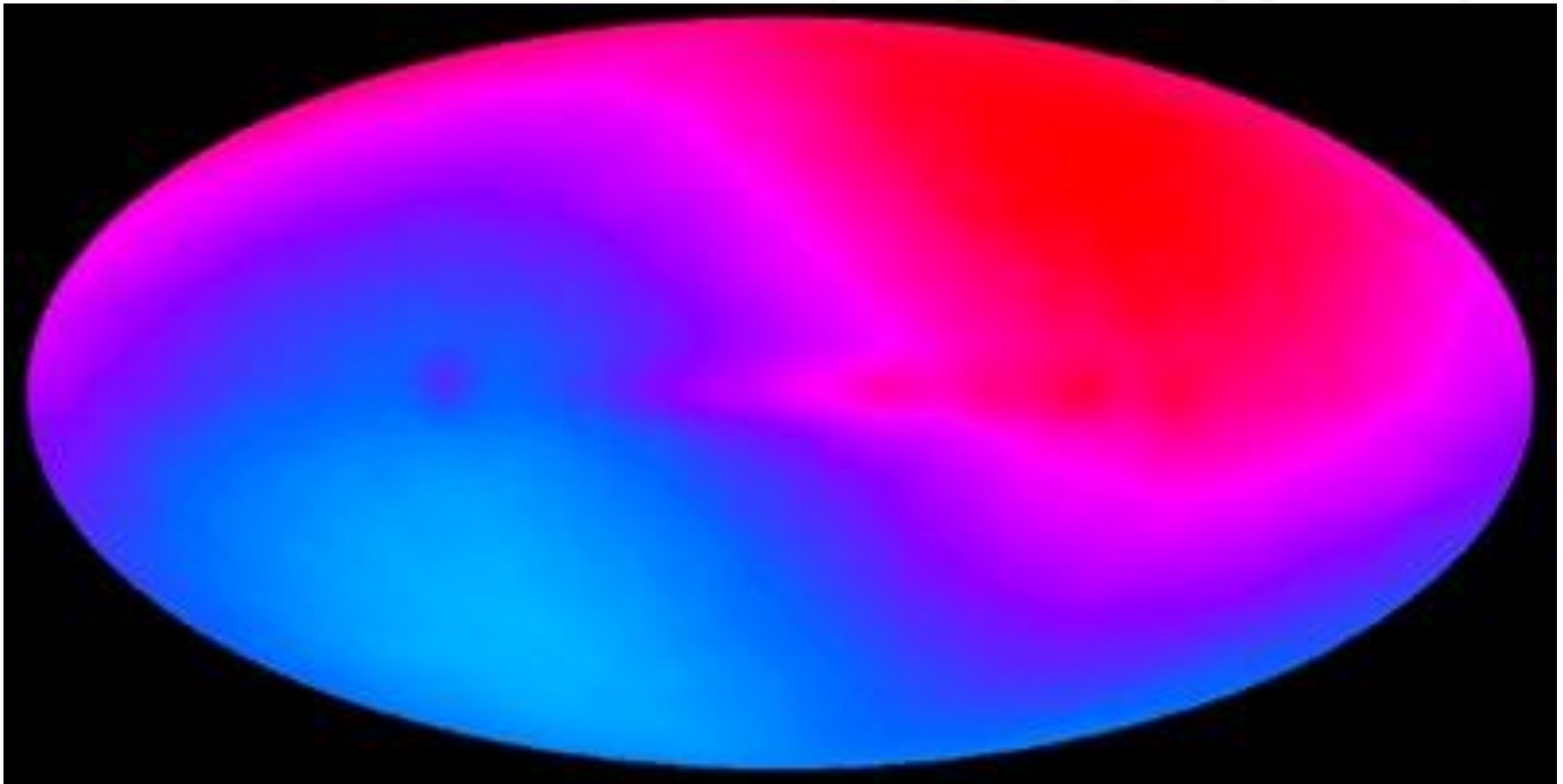
YADBOUD UNIVER

# Anisotropie in CMB



**Temperatuur van de CMB (Cosmic Microwave Background) aan de hemel in kleurcode: 0 K is blauw en 4 K is rood (COBE). Het melkwegvlak loopt horizontaal en het centrum van de figuur is in de richting van het centrum van ons Melkwegstelsel.**

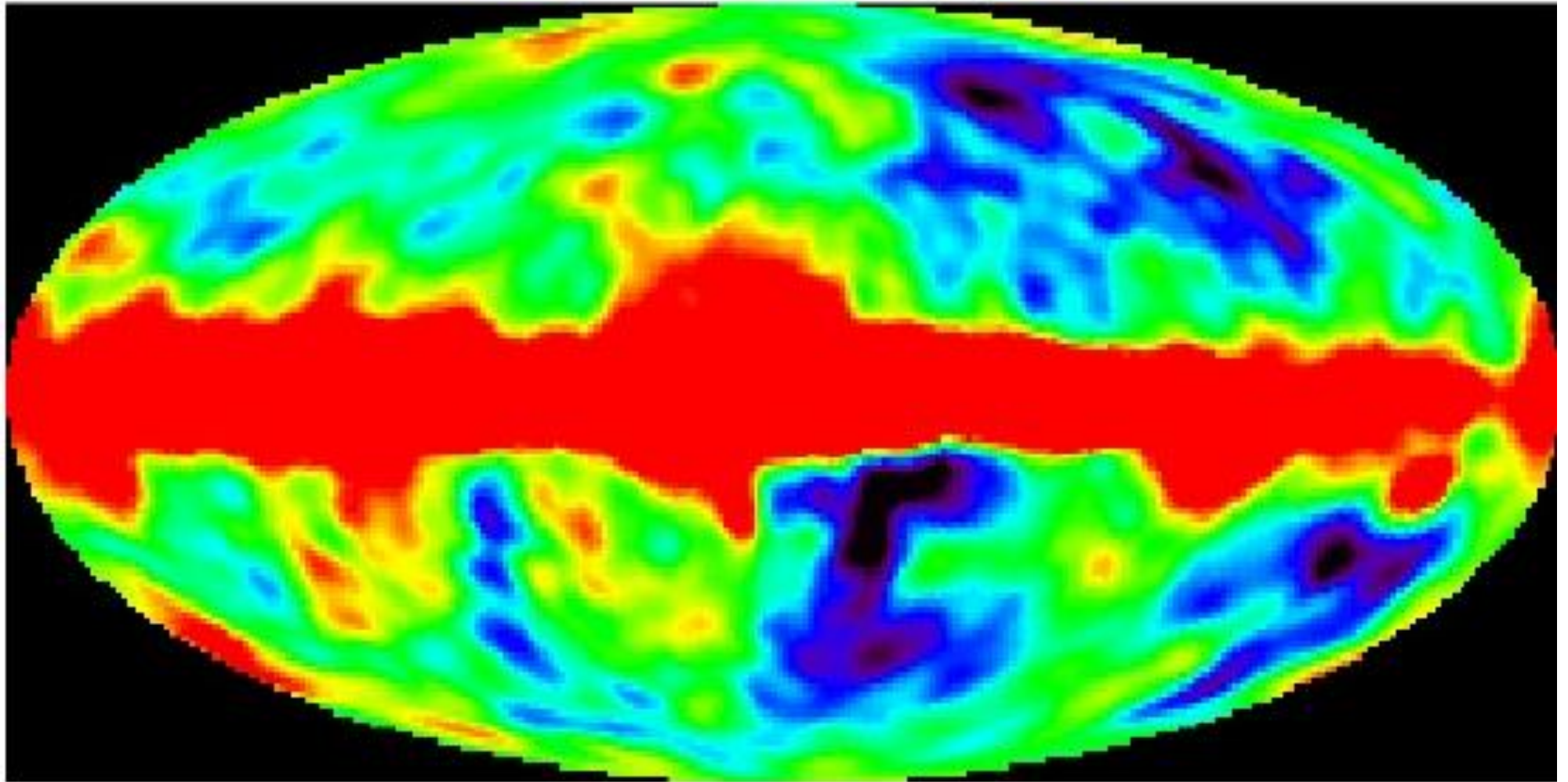
Dipoolanisotropie:  $\Delta T = \pm 0,00335 \text{ K}$



Fluctuaties in de CMB-temperatuur aan de hemel op een schaal tussen 2,724 K (blauw) en 2.732 K (rood) met COBE.

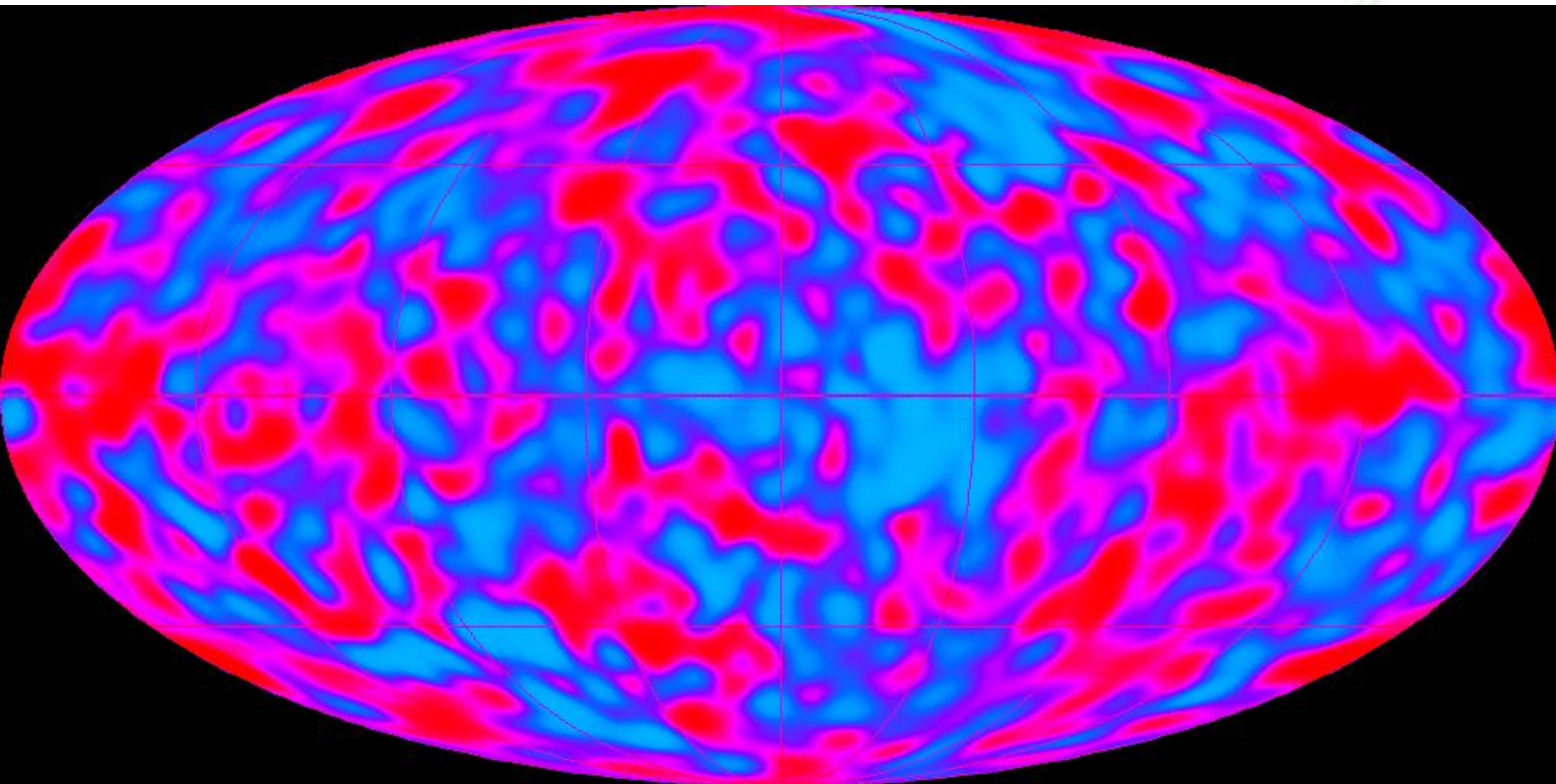


Temperatuurfluctuaties na aftrekken van de dipoolcomponent:  
van blauw naar rood  $\Delta T = 0,000018$  K

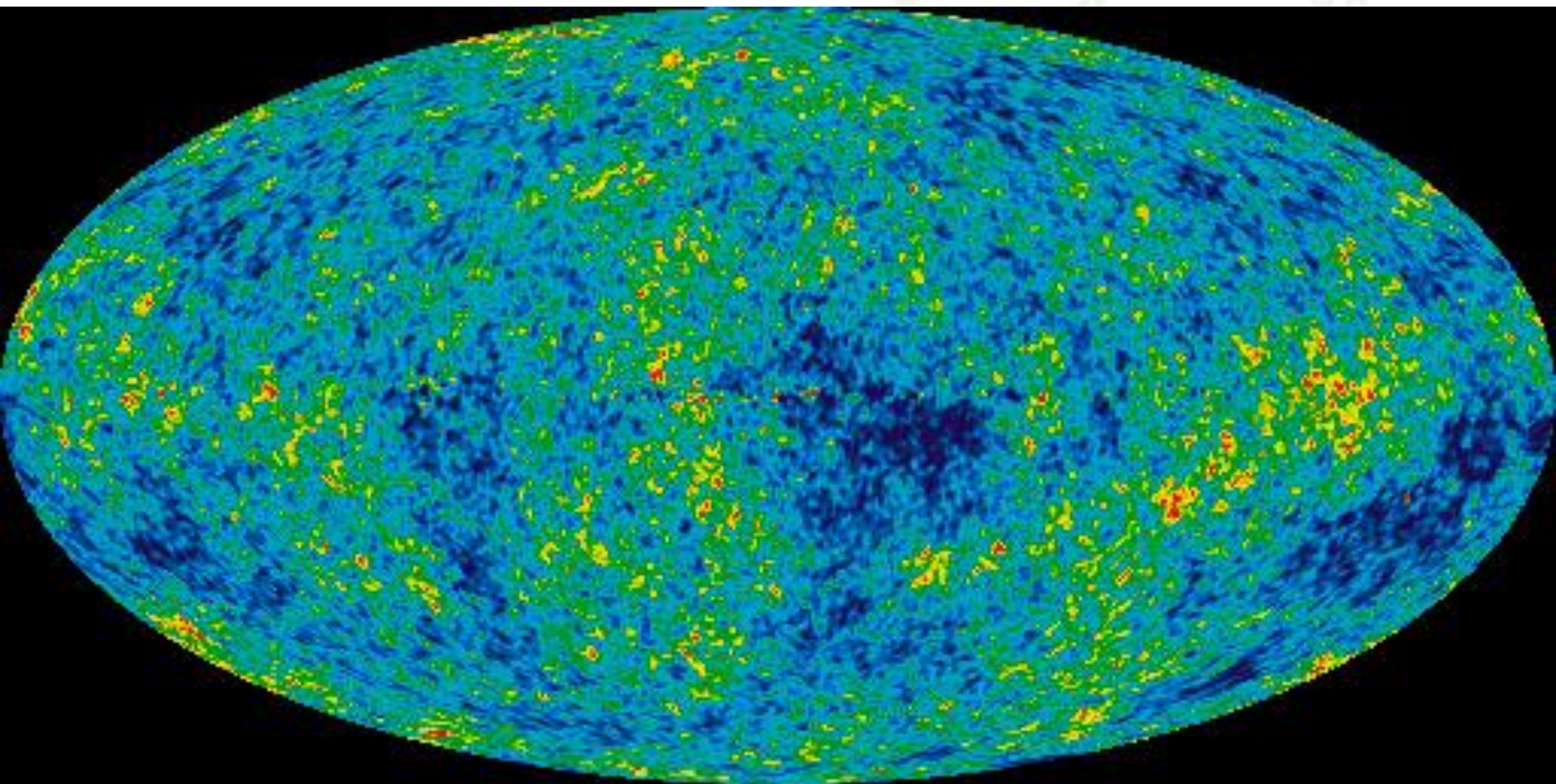


# Fluctuaties in achtergrondstraling!

Na aftrek van dipoolcomponent en galactische straling van stof (termisch), heet gas (vrij-vrij),

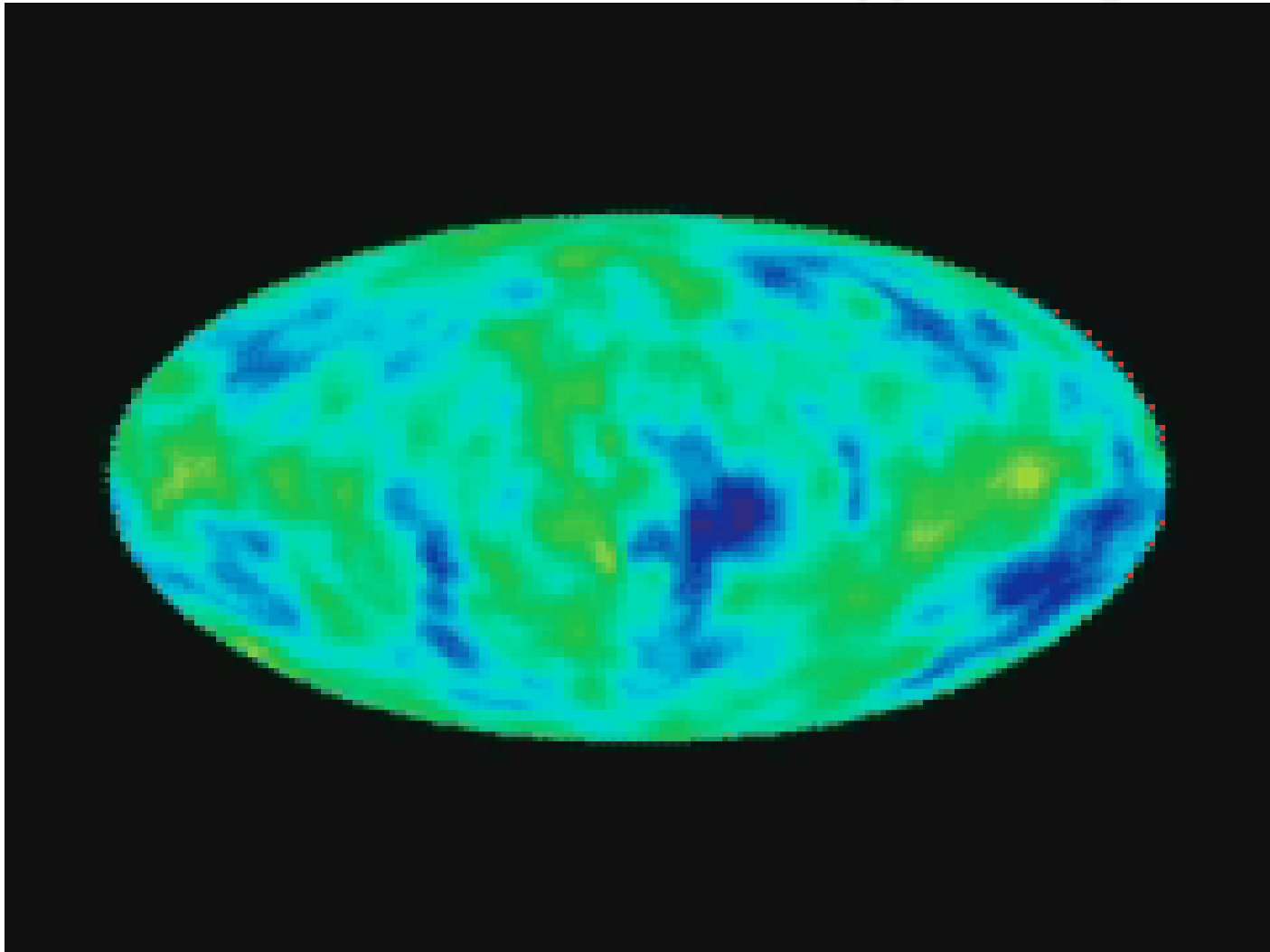




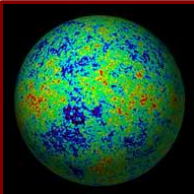


WMAP



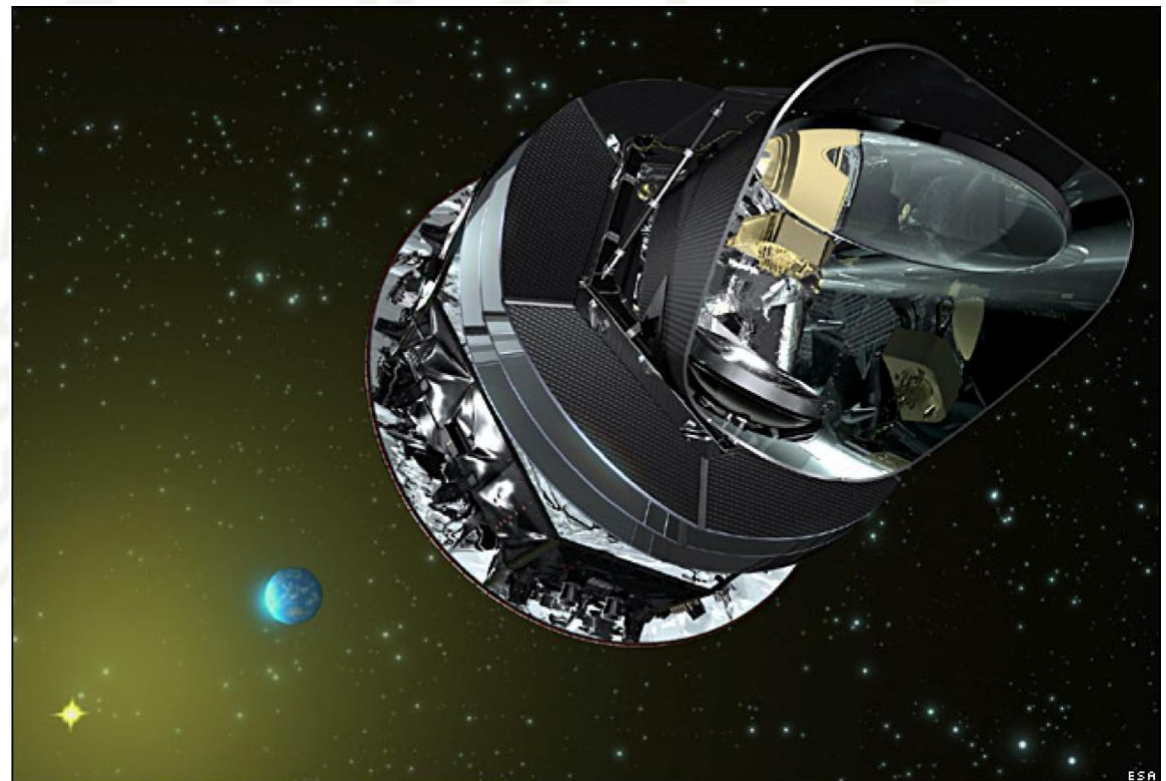


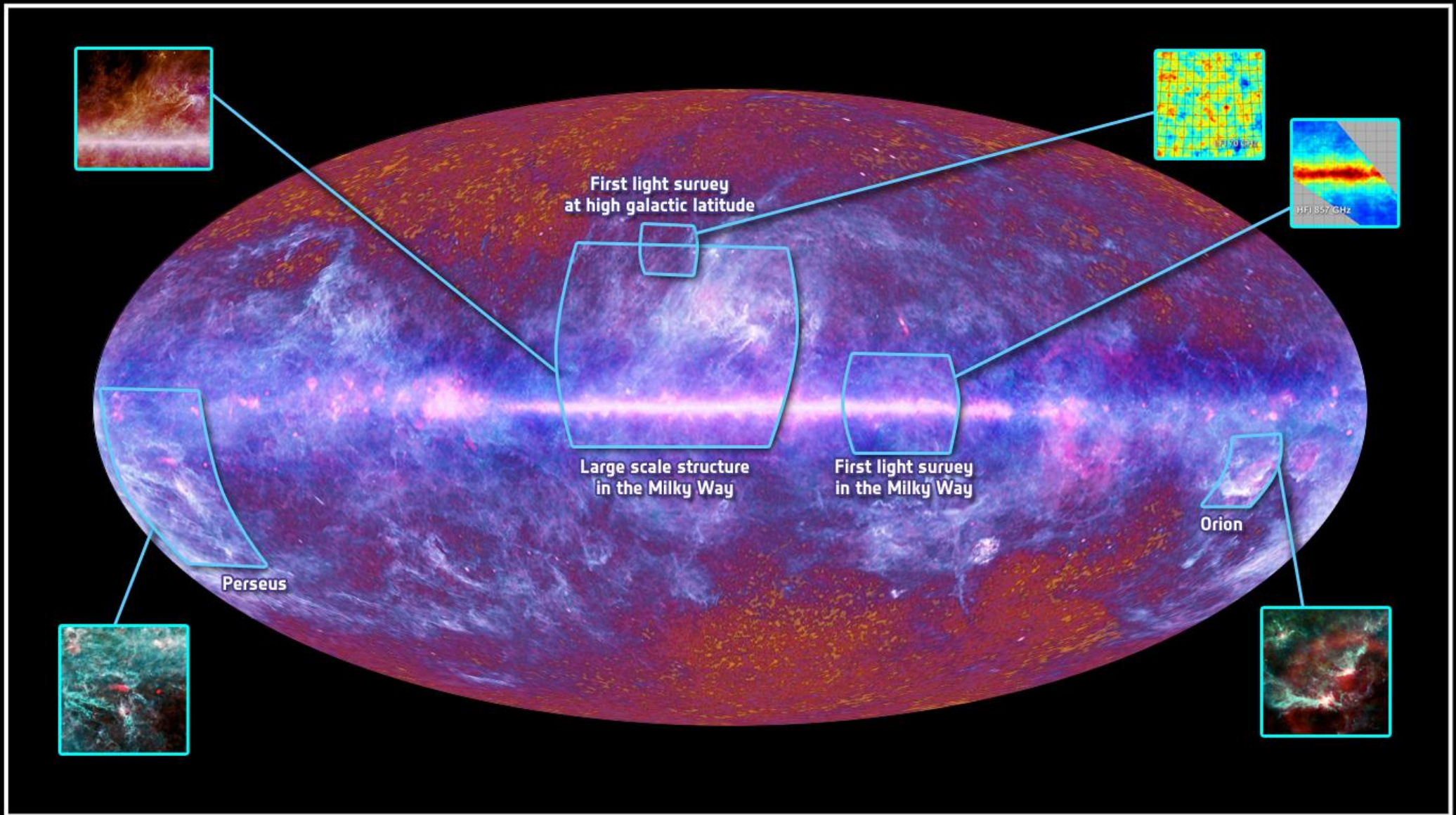
OF ASTROPHYSICS  
NIJMEGEN  
Radboud University  
Radboud University



# Planck

- Nieuwe ESA satelliet
- Gelanceerd in 2009 (samen met Herschel)
- Gaat nog gedetailleerdere kaart van de achtergrondstraling maken

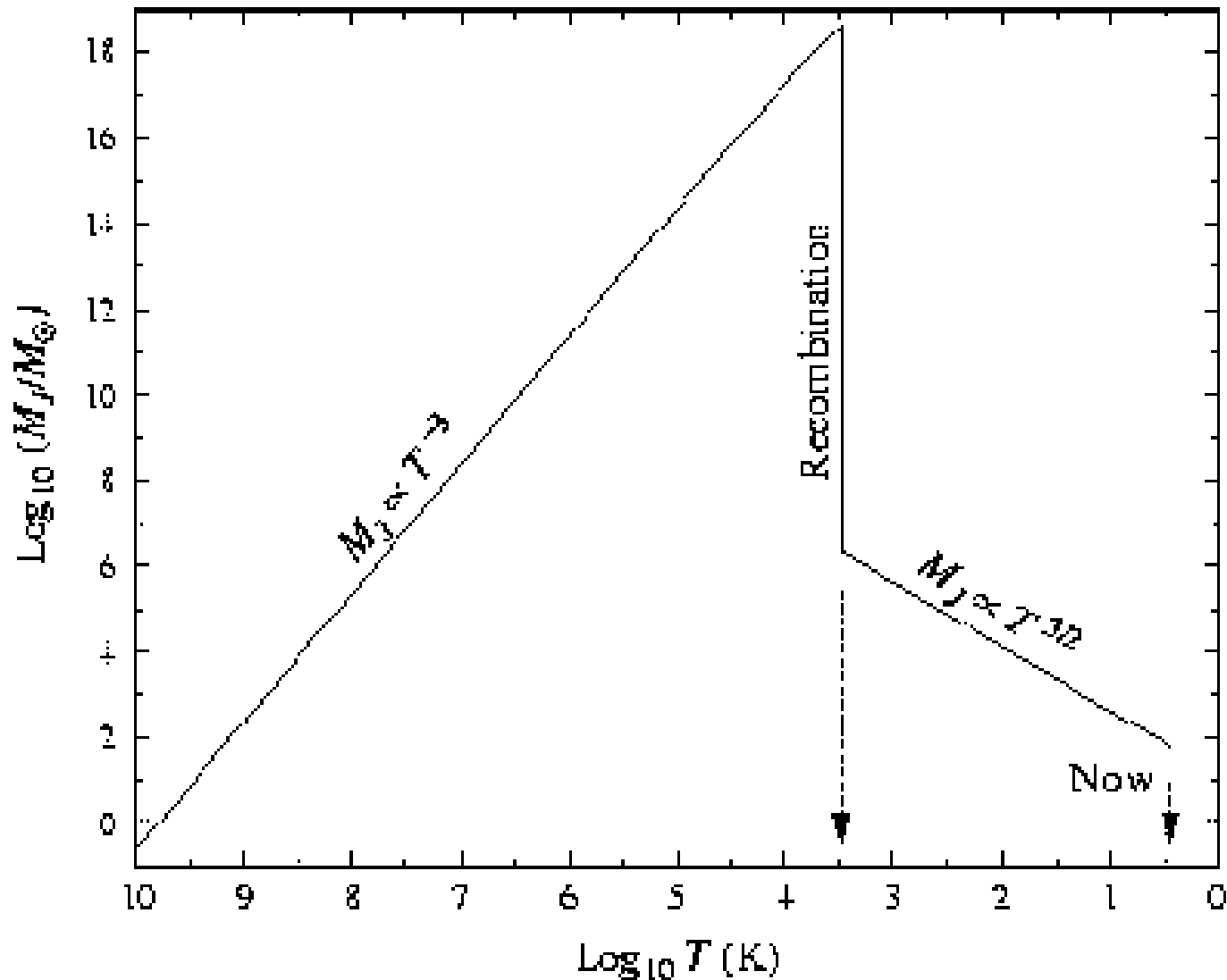






# Jeansmassa

Een gaswolk kan zich slechts samentrekken als zijn interne gasdrukkracht klein genoeg is ten opzichte van de zwaartekracht:



# WMAP

structuurvorming uit  
fluctuaties

Ook: meting uitdijing!  
(laatste college)

